

The SANAP Cosmic-Ray Neutron Monitor Programme

Harm Moraal

**North-West University
Potchefstroom**

Our group

1. Harm Moraal - PI
2. Helena Krüger - Co-I
3. Gert Benadé – electronics engineer
4. Anne Mans – data and station manager
5. Godfrey Mosotho – M.Sc. student
6. Renier Fuchs – M.Eng. student
7. Ruan Nel – M.Eng. student
8. Henrdik Krüger – expedition member SANAE

Pieter Stoker – Emeritus and founder

Cosmic Rays

- Charged **particles** - 90% protons, 5% He nuclei, 3% heavier atomic nuclei, 2% electrons
- Characterised by very high energies (10^6 - 10^{20} eV)

One particle = cricket ball at 140 km per hour.

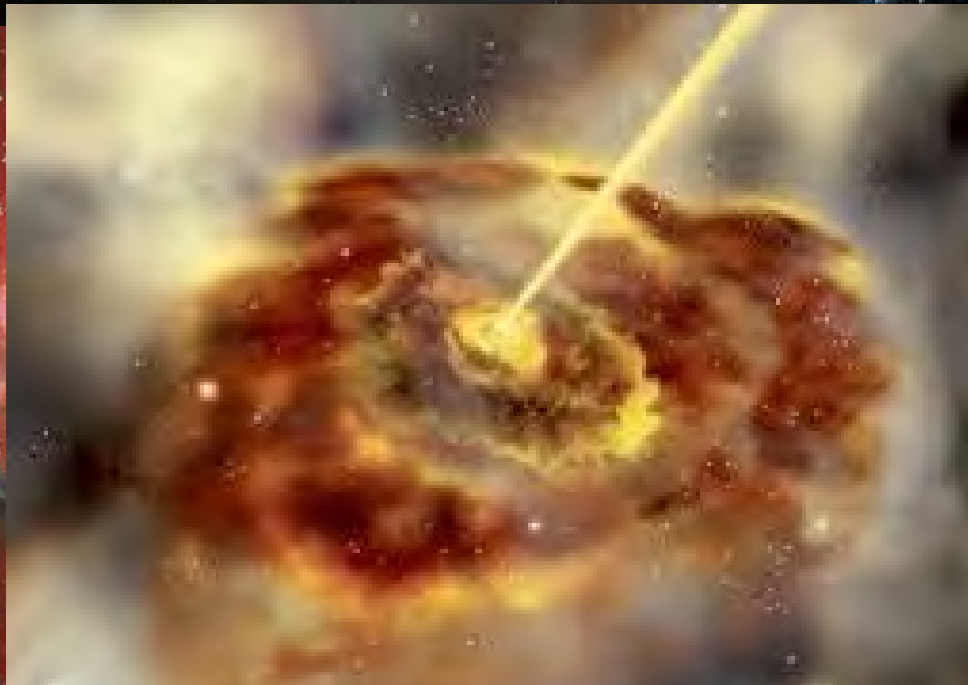
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Come from the Cosmos



Particle vs. Photon Astronomy

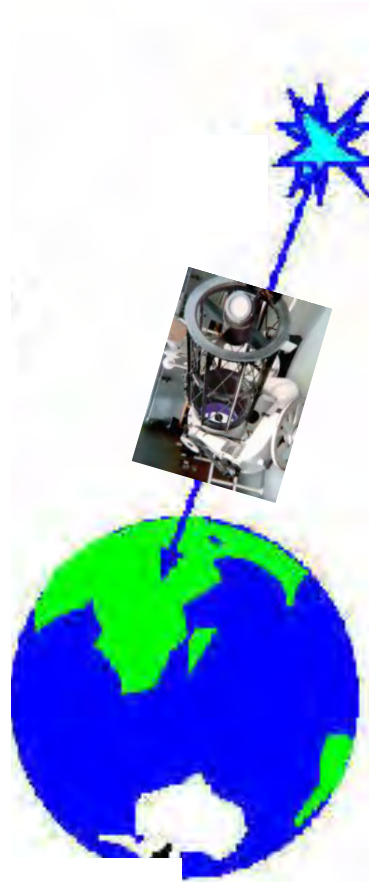


Photons

1. Where
2. How bright
3. Colour



Particle vs. Photon Astronomy



Photons

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Particle vs. Photon Astronomy

Particles:
No such information
Because of magnetic fields

.....like a bead on an
elastic band



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Particle vs. Photon Astronomy



Photons

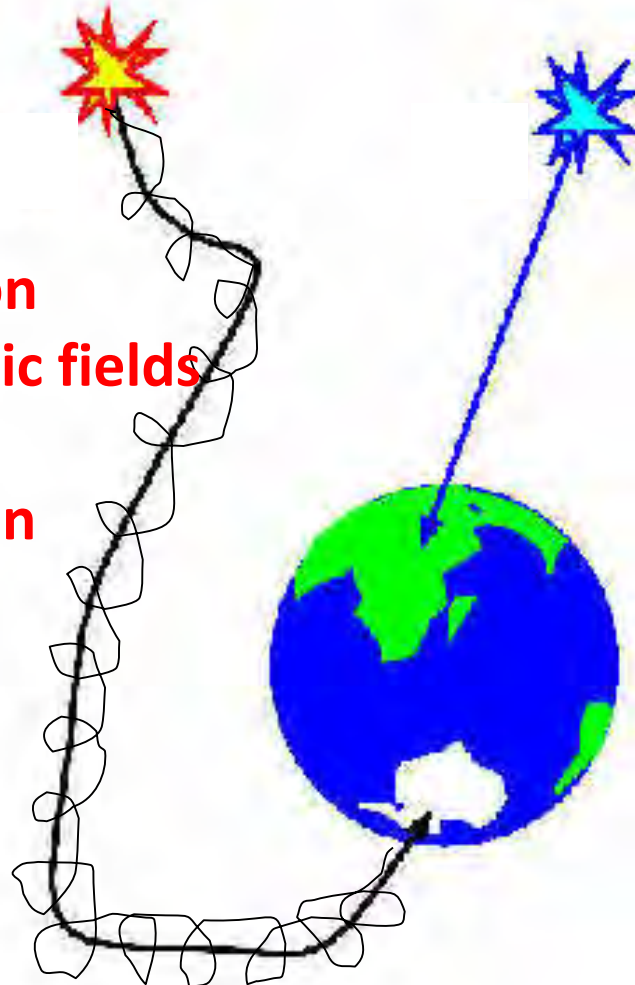
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Photons

1. Where
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Victor Hess, 7 August 1912



Victor Hess, 7 August 1912



100 years later (+2)

Cosmic Rays in the Heliosphere

Harm Moraal

North-West University

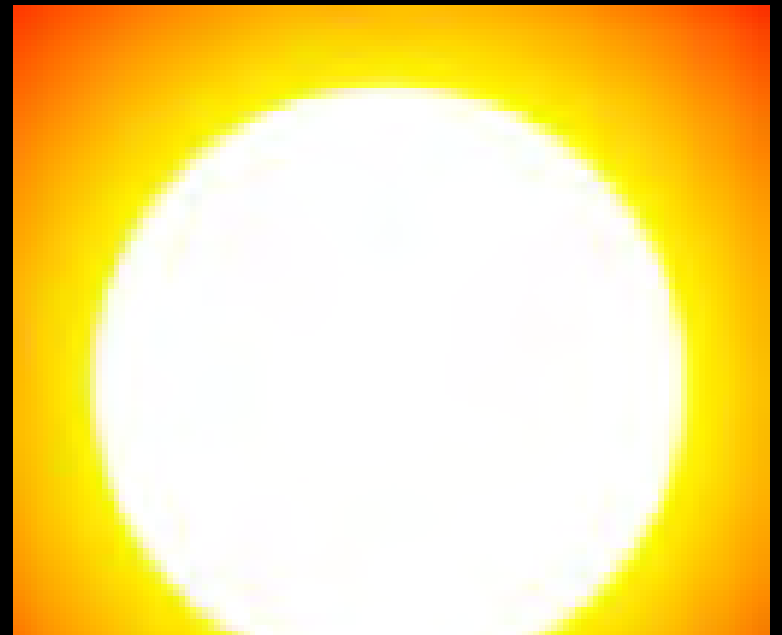
Potchefstroom, South Africa

Bad Saarow, Germany, 8 August 2012

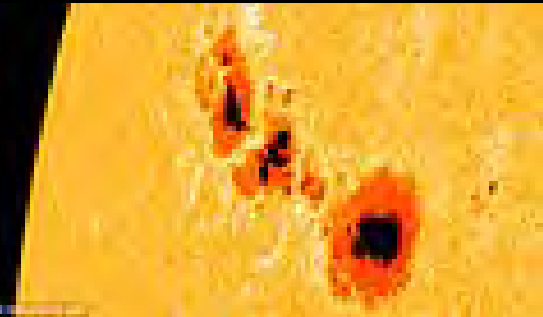
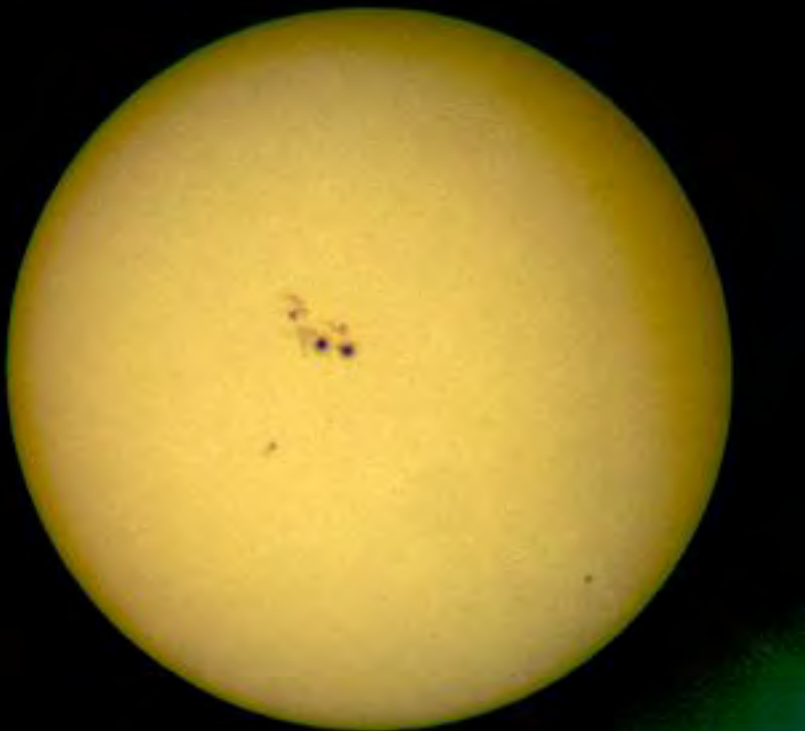
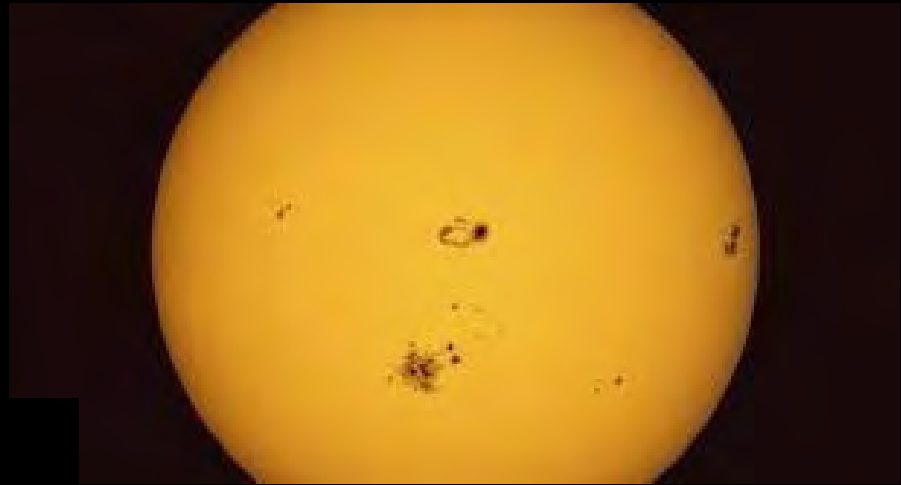
Bad Saarow Railway Station



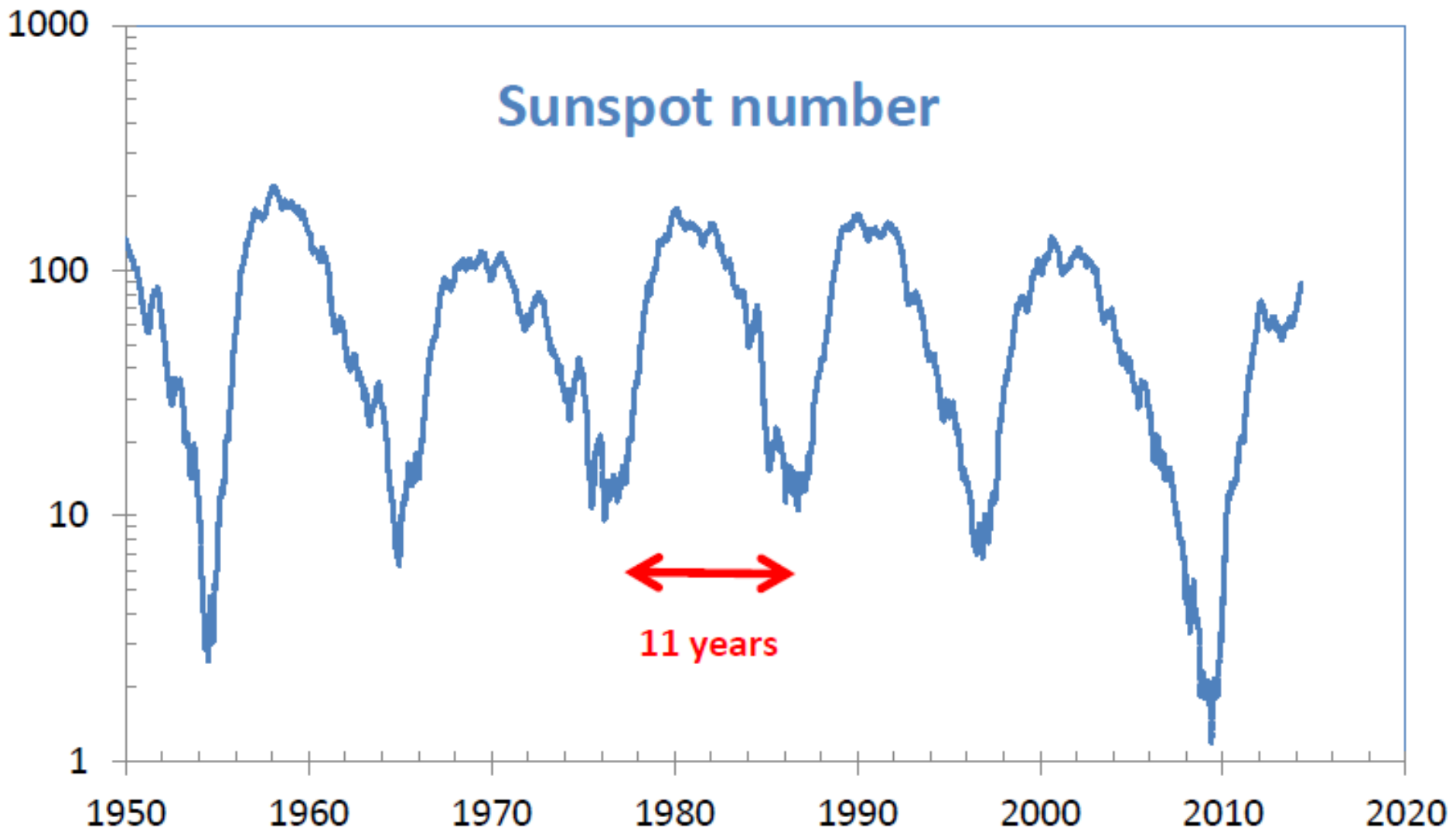
The Sun



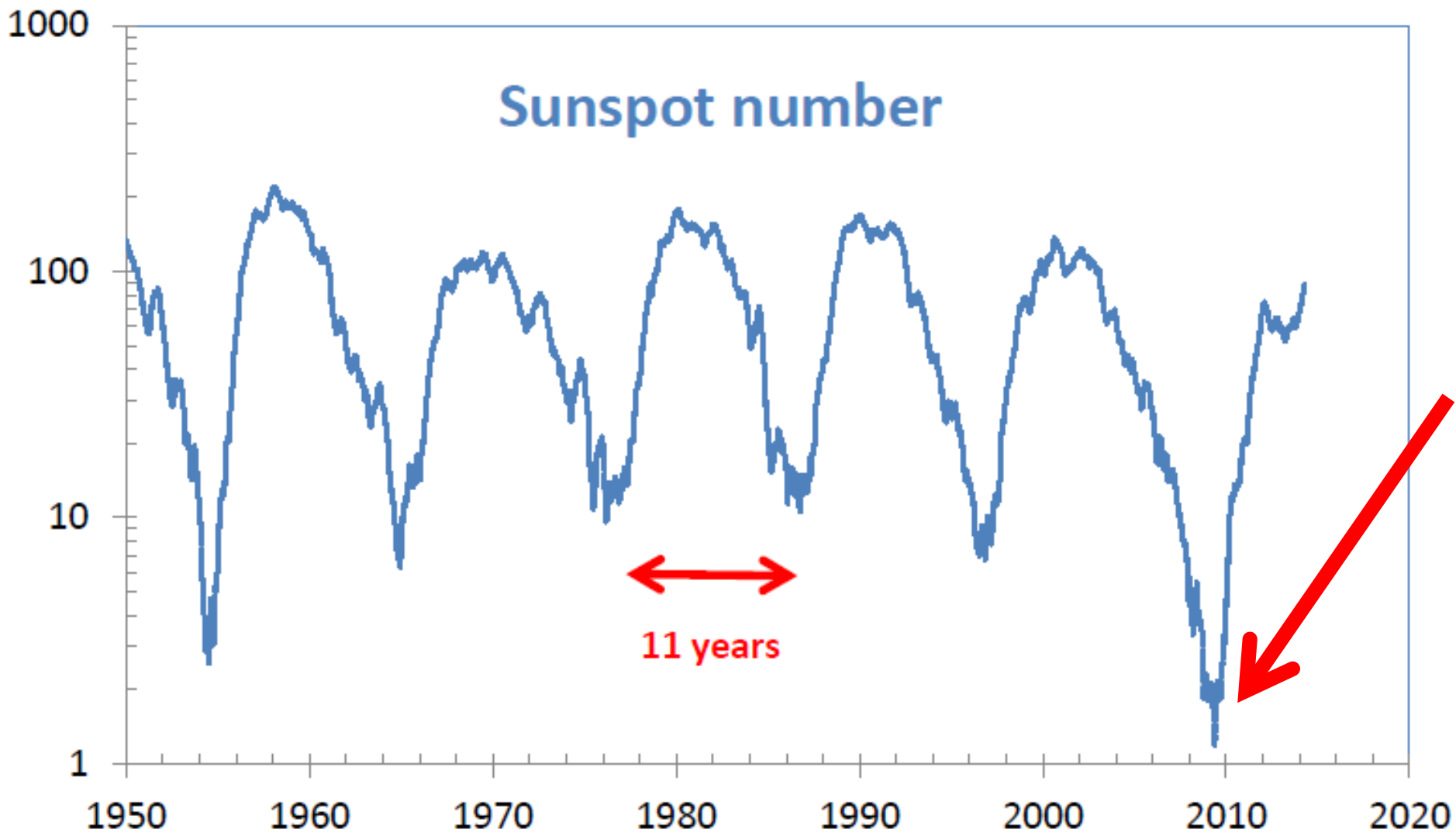
Sunspots



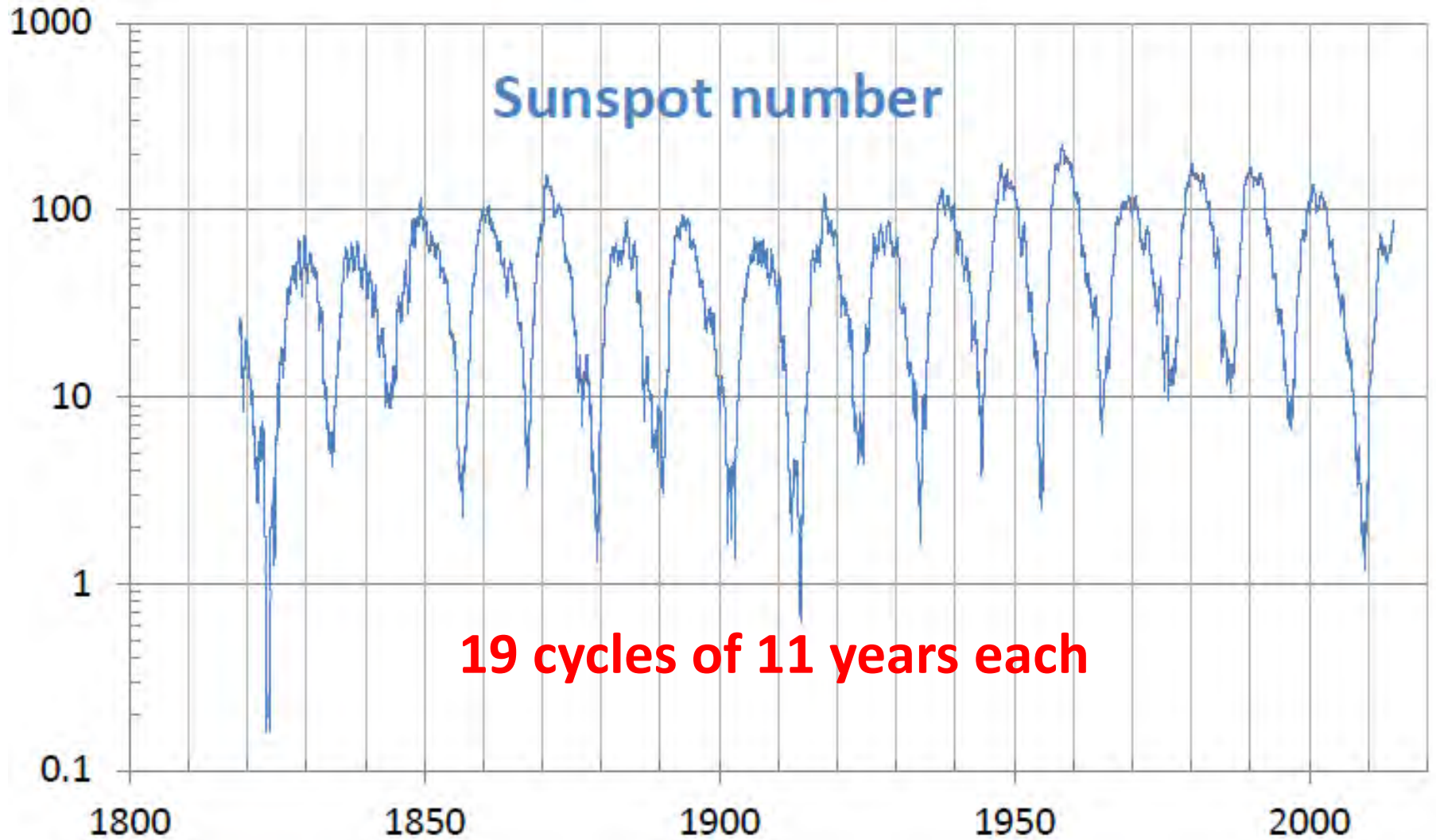
Sunspots



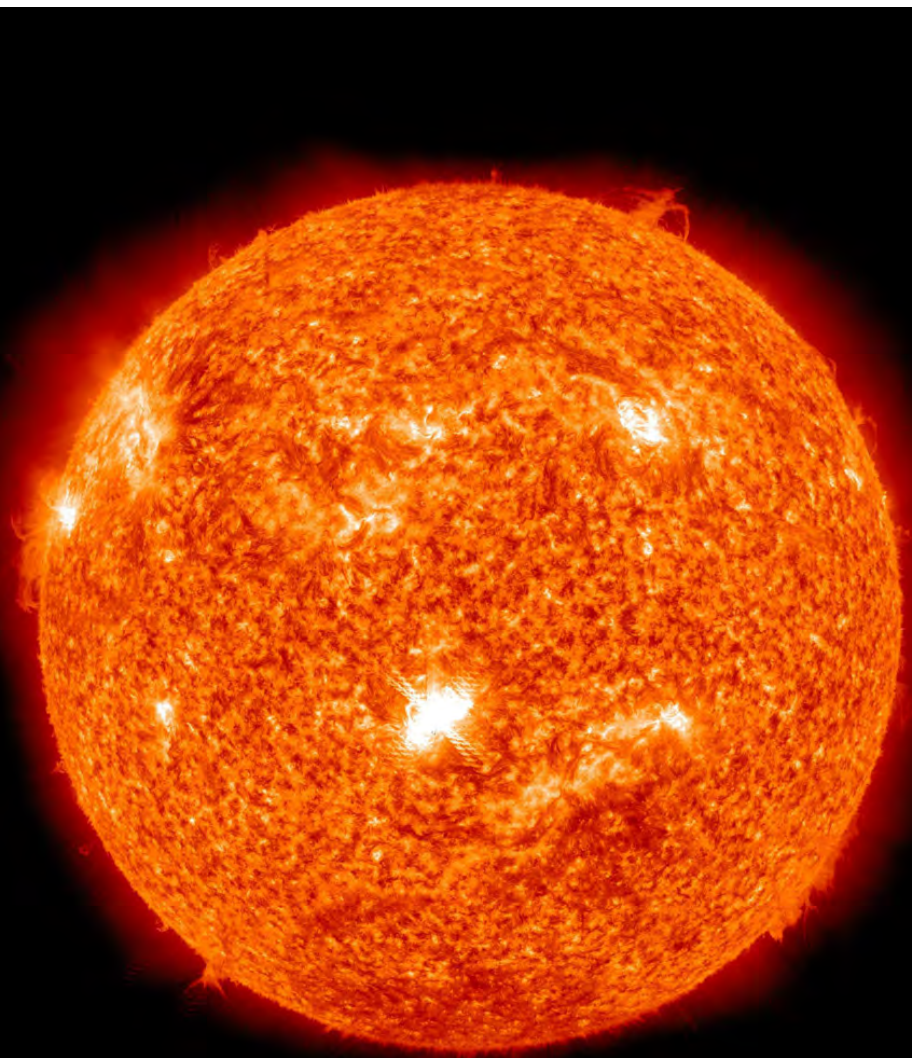
Sunspots



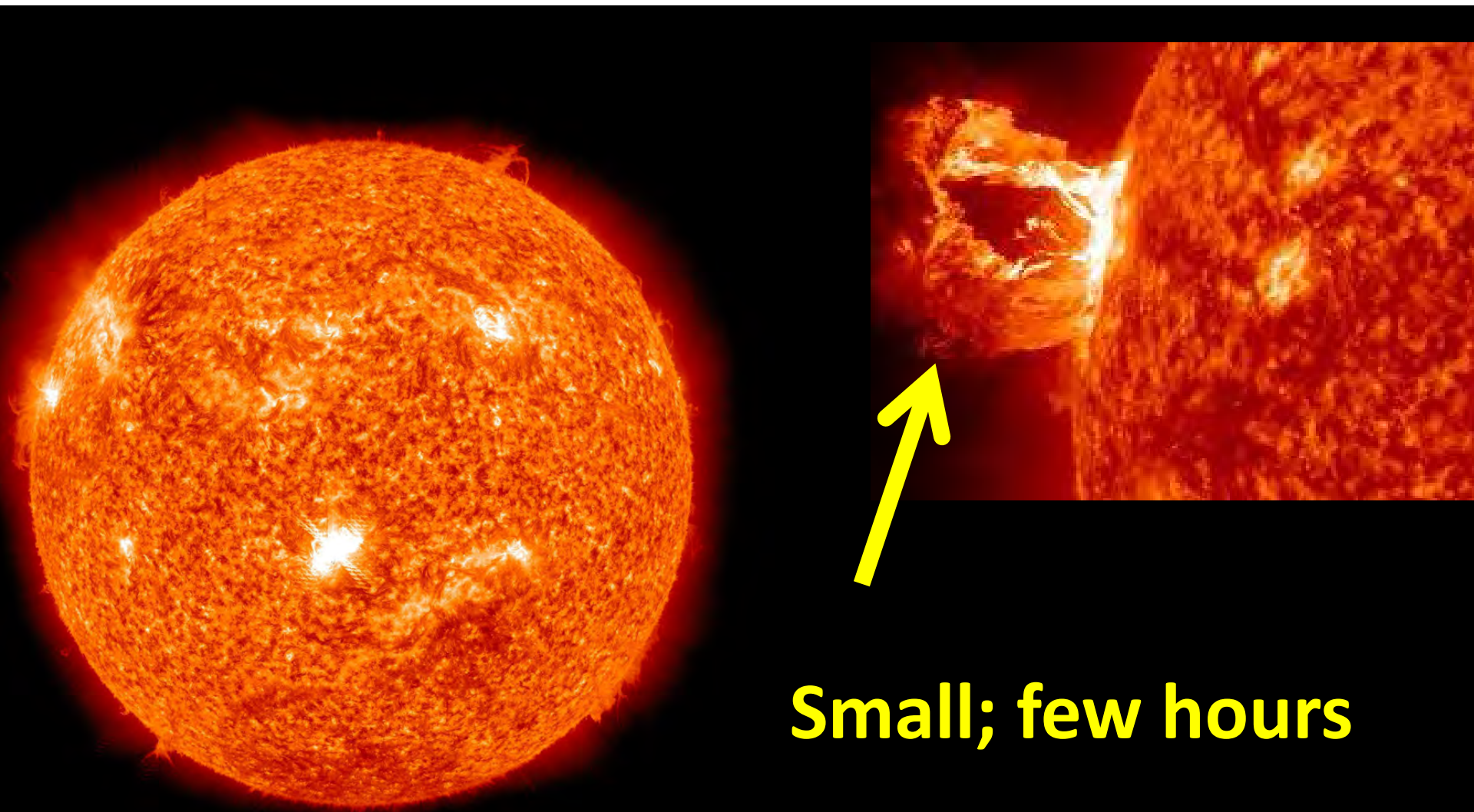
Sunspots since 1818



Solar Flare

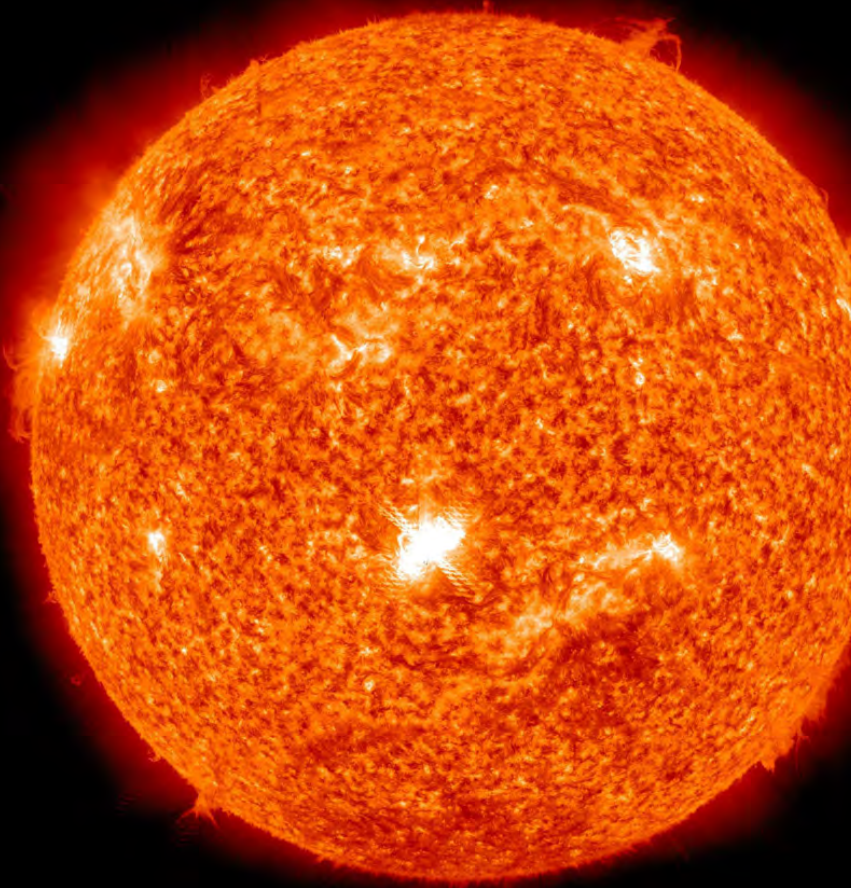


Solar Flare

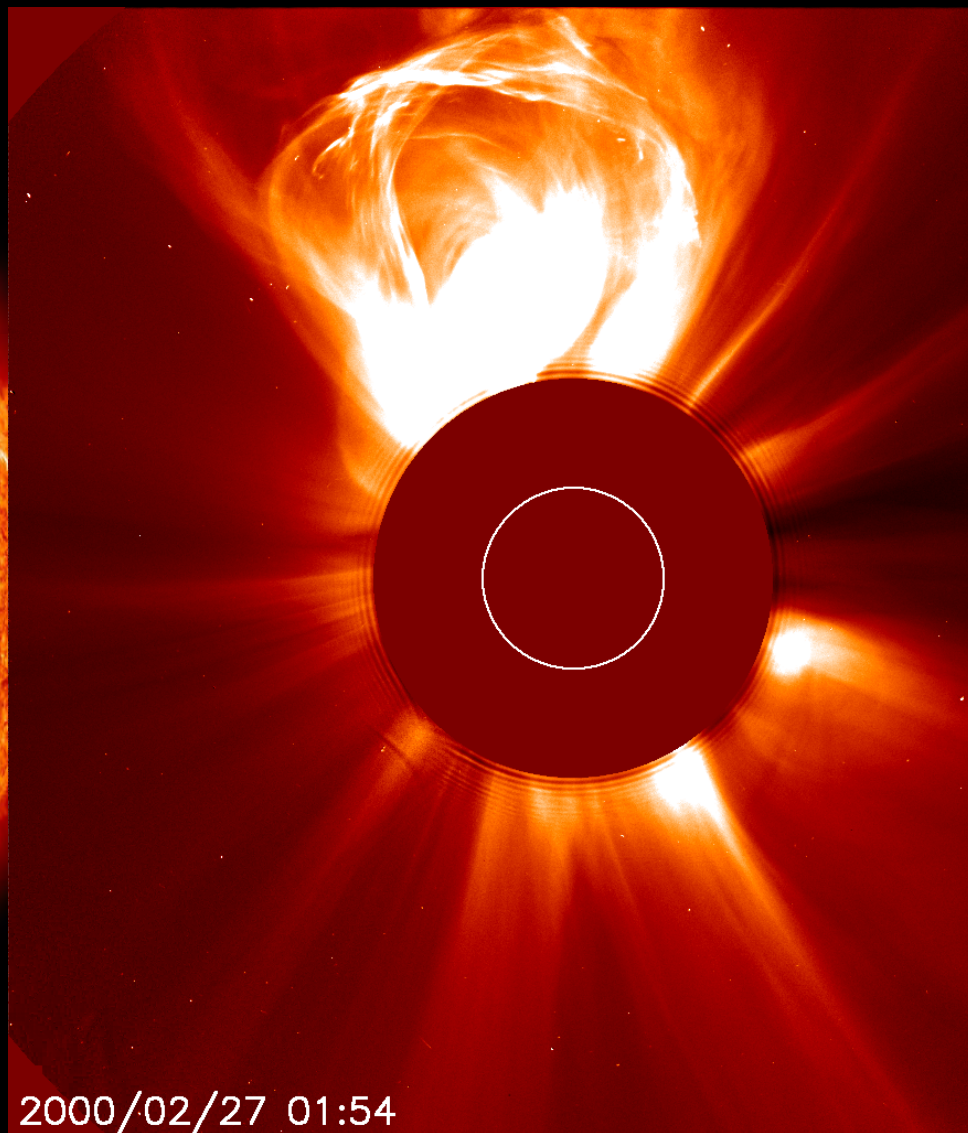


Small; few hours

Coronal Mass Ejection



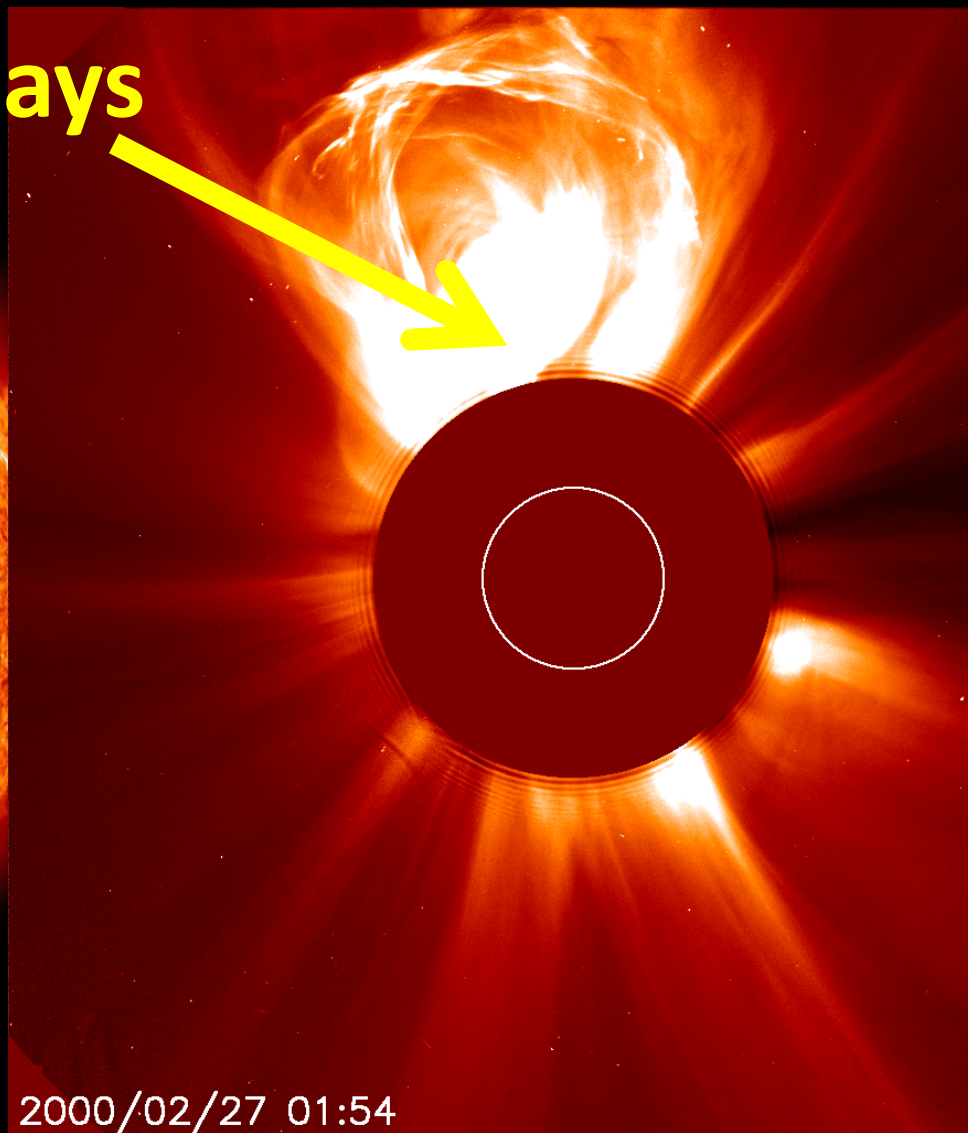
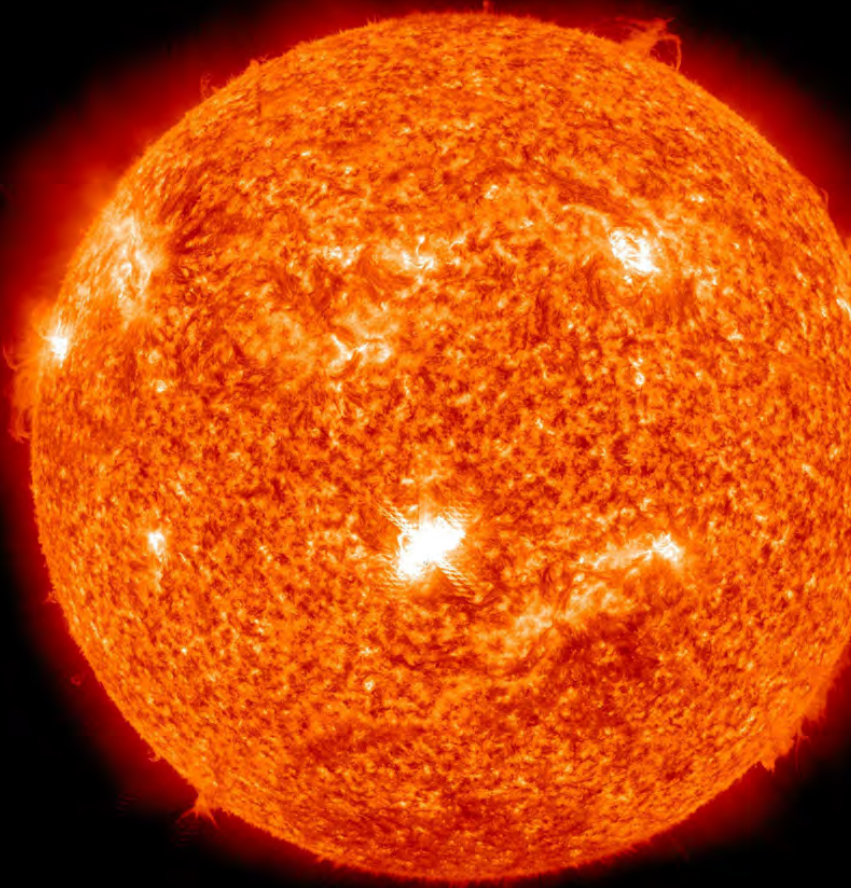
SDO/AIA 304 2011-02-13 17:36:45 UT



2000/02/27 01:54

Coronal Mass Ejection

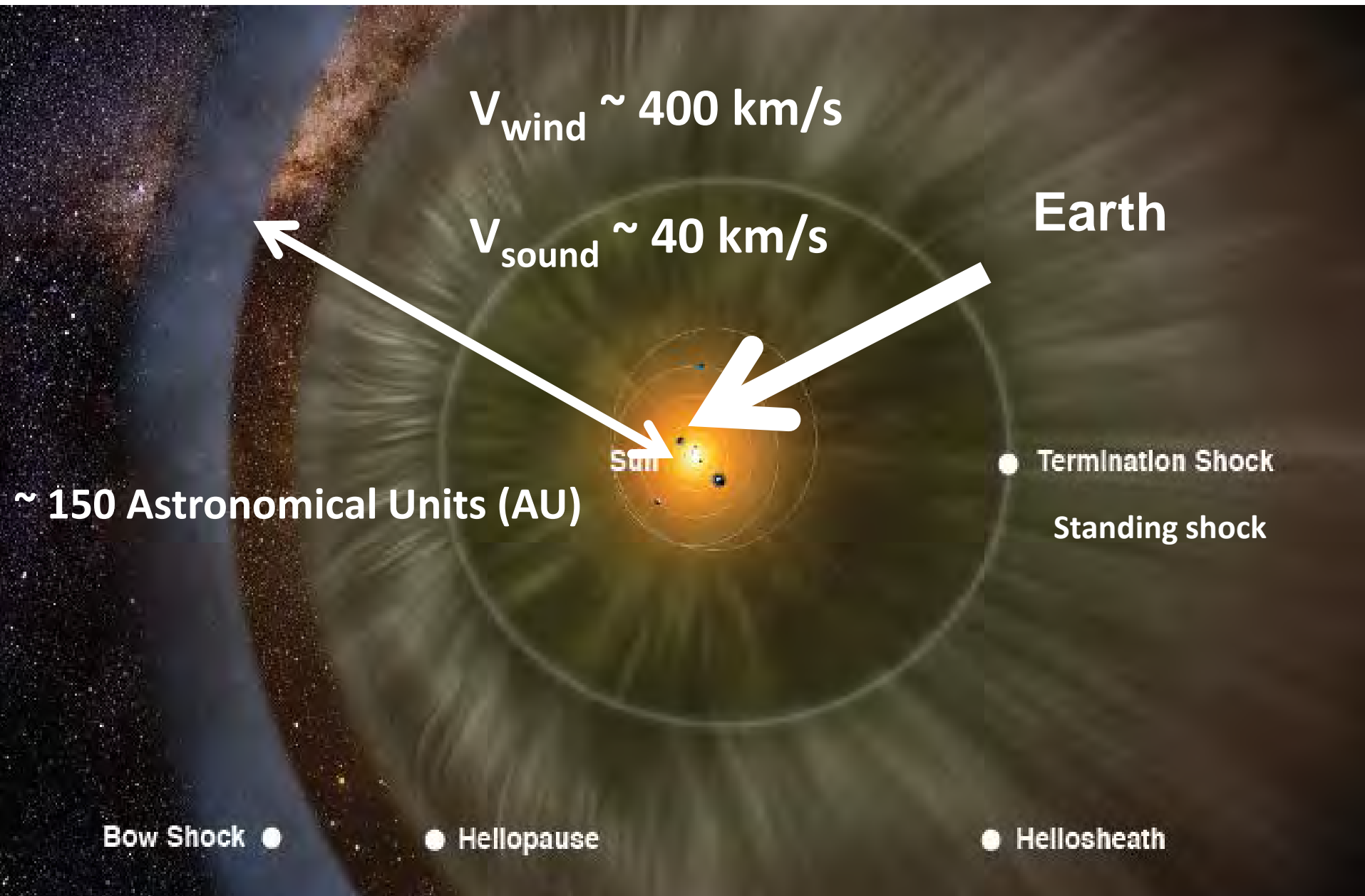
Big; days



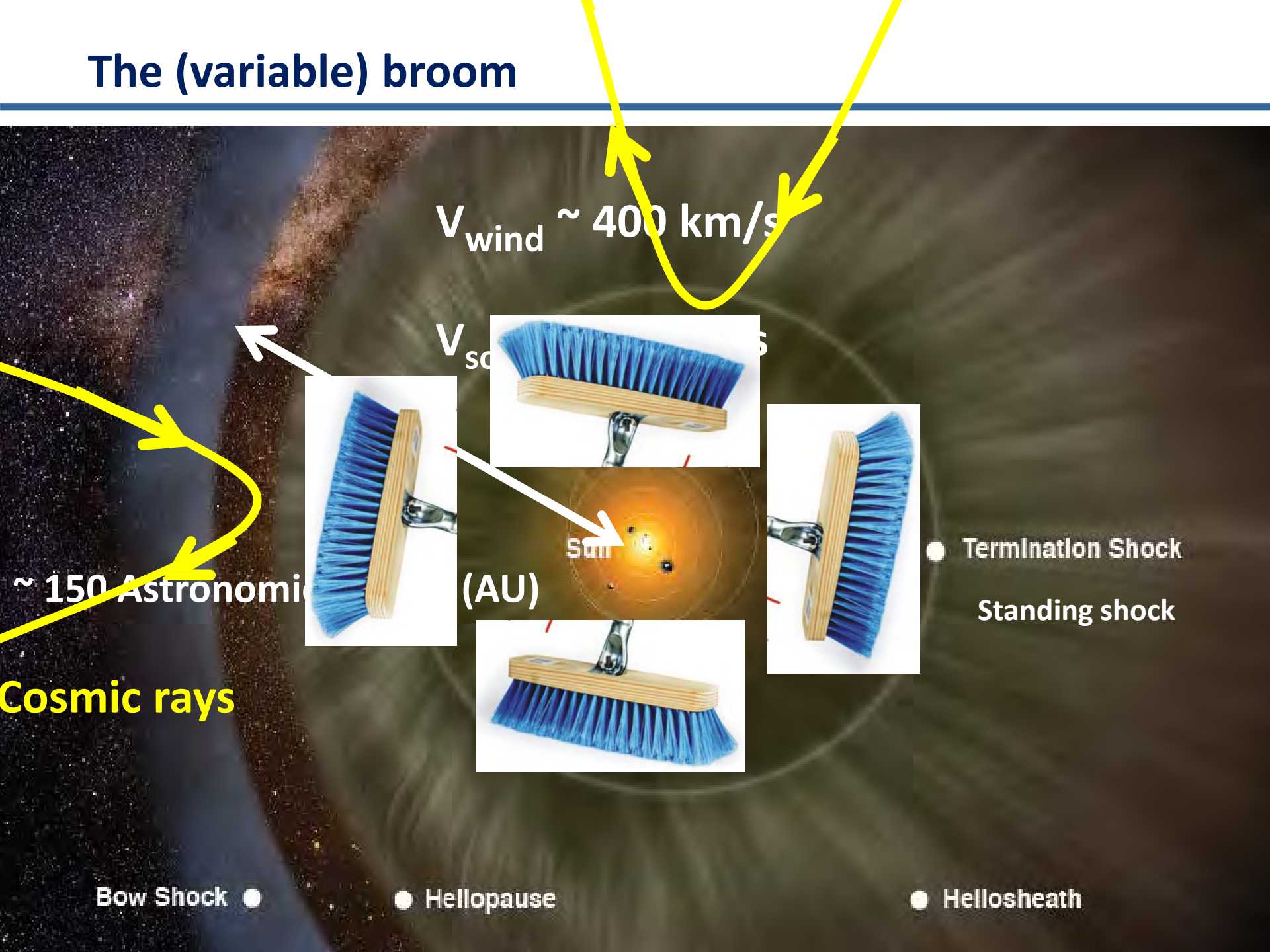
SDO/AIA 304 2011-02-13 17:36:45 UT

2000/02/27 01:54

The Solar Wind and Heliosphere

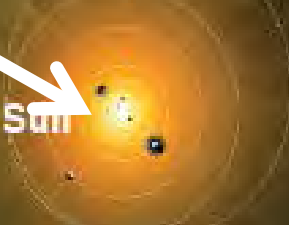


The (variable) broom



$V_{wind} \sim 400 \text{ km/s}$

V_{sc}



● Termination Shock
Standing shock

● Bow Shock

● Helopause

● Heliosheath

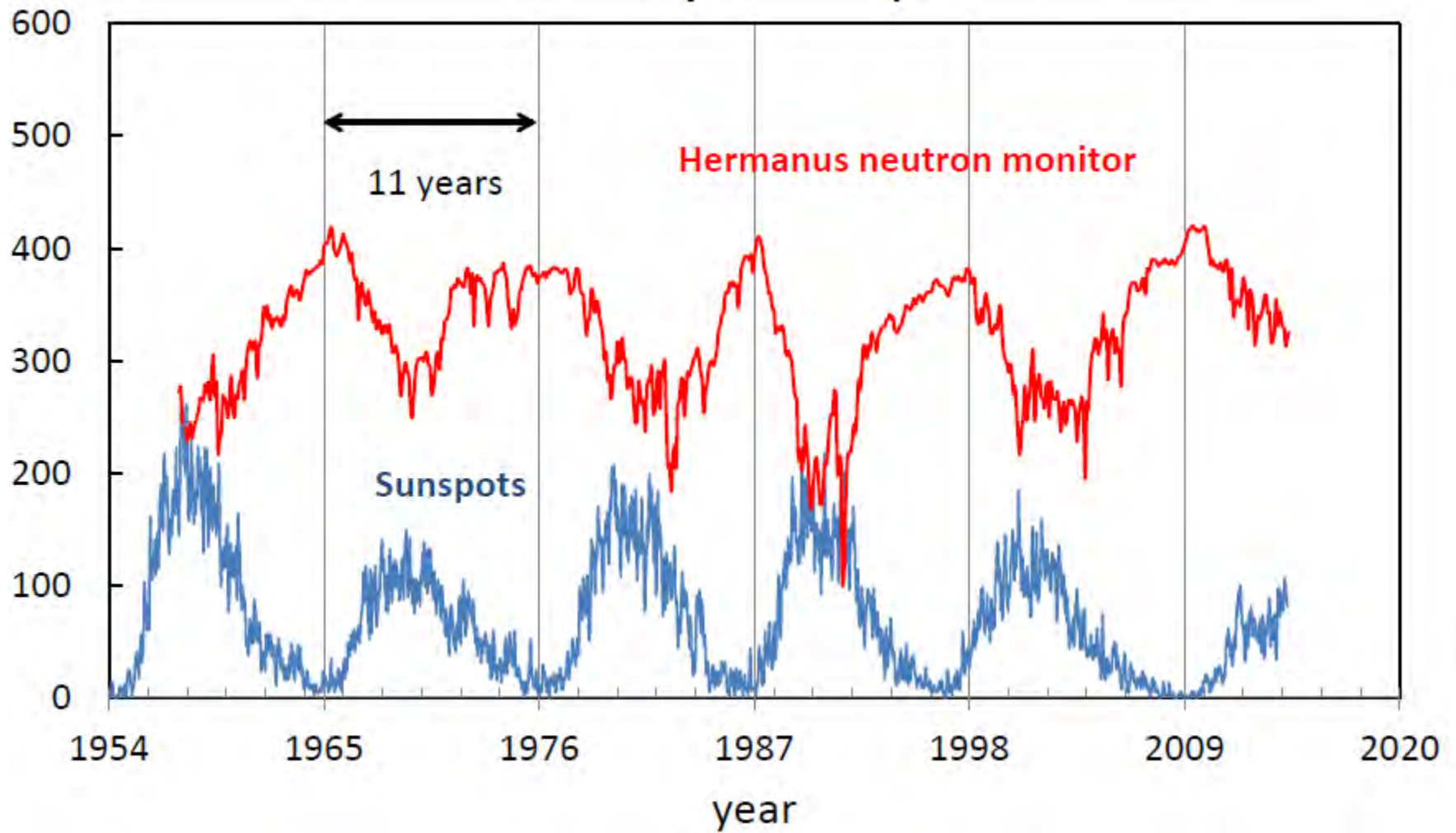
$\sim 150 \text{ Astronomical Units}$

Cosmic rays

Sun

(AU)

Cosmic rays and sunspots

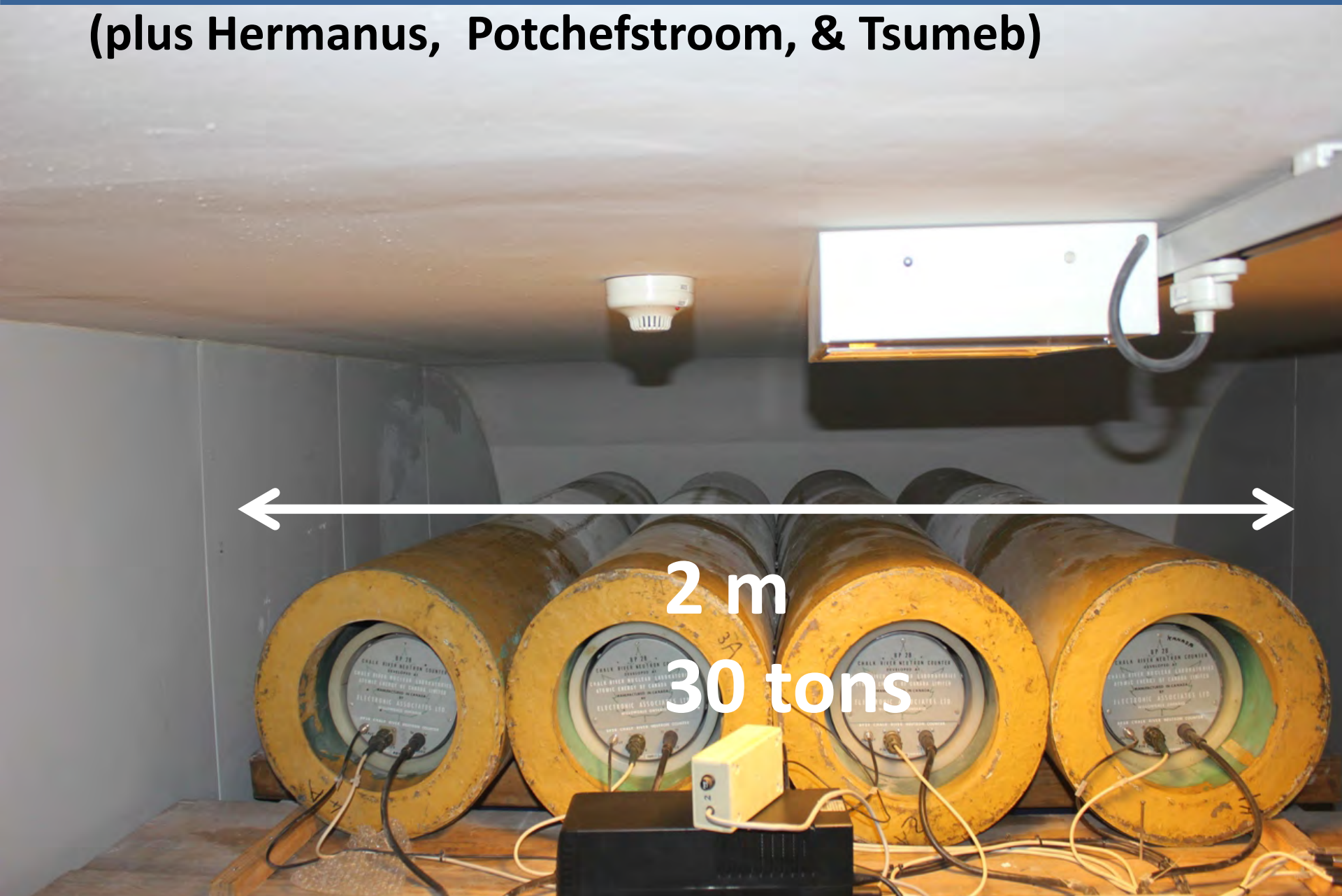


Sanae Neutron Monitor



Sanae Neutron Monitor

(plus Hermanus, Potchefstroom, & Tsumeb)



2 m

30 tons

Mini neutron monitors



Polarstern



Mini neutron monitors

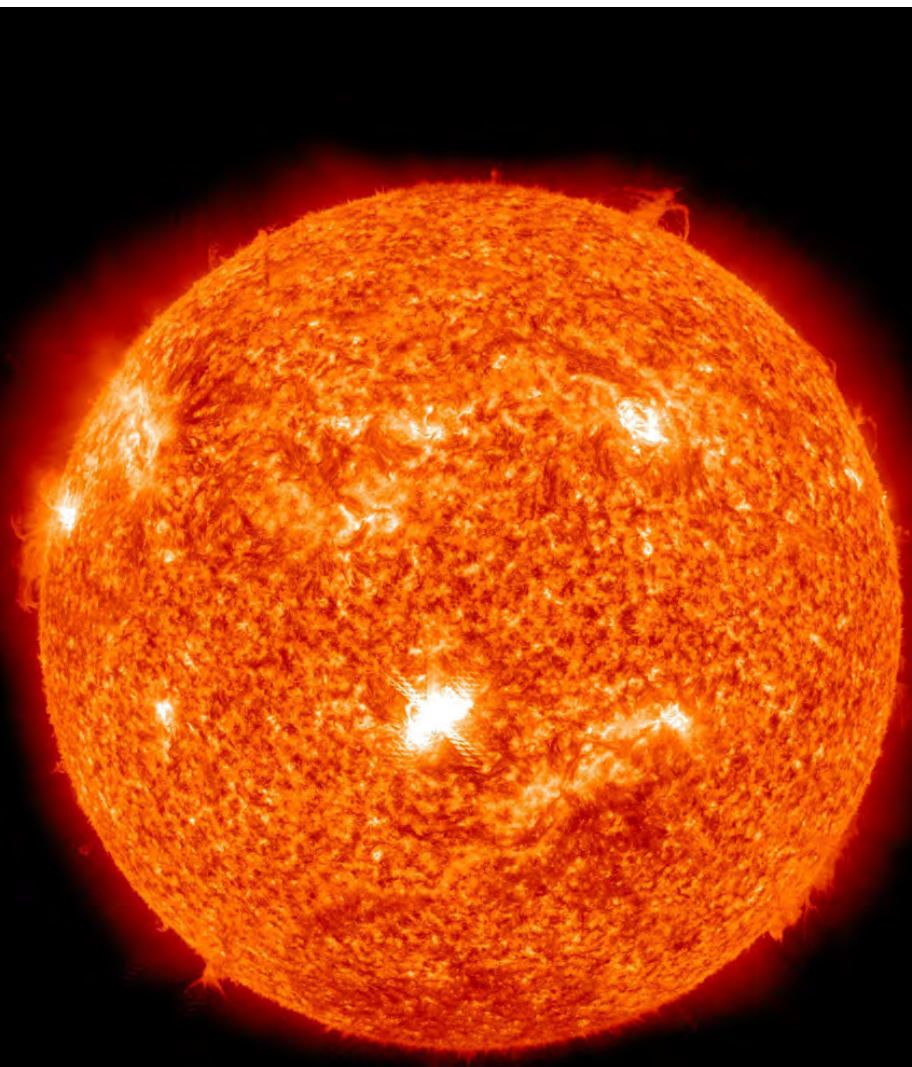


Polarstern

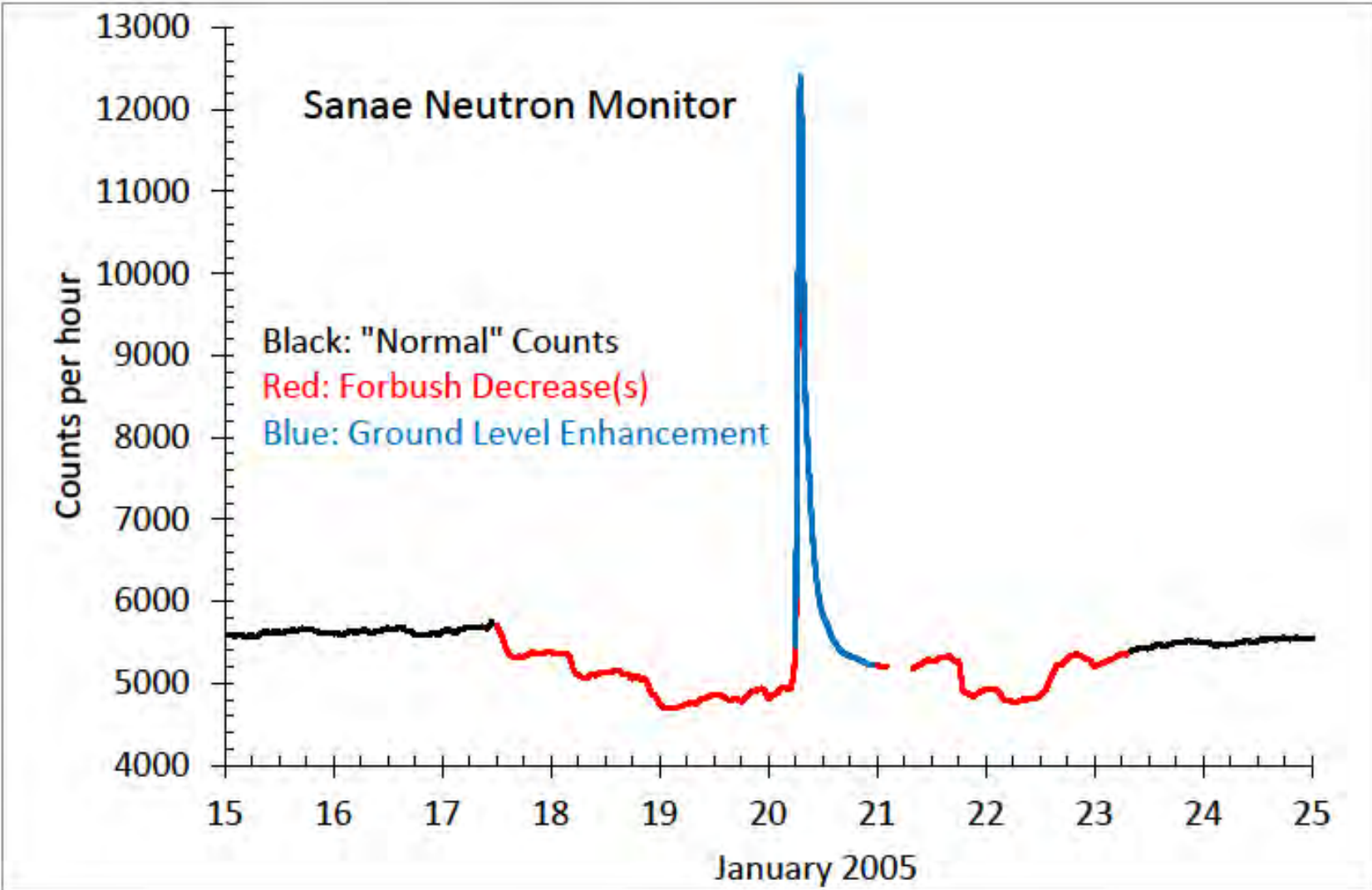


Posters
Gert Benadé

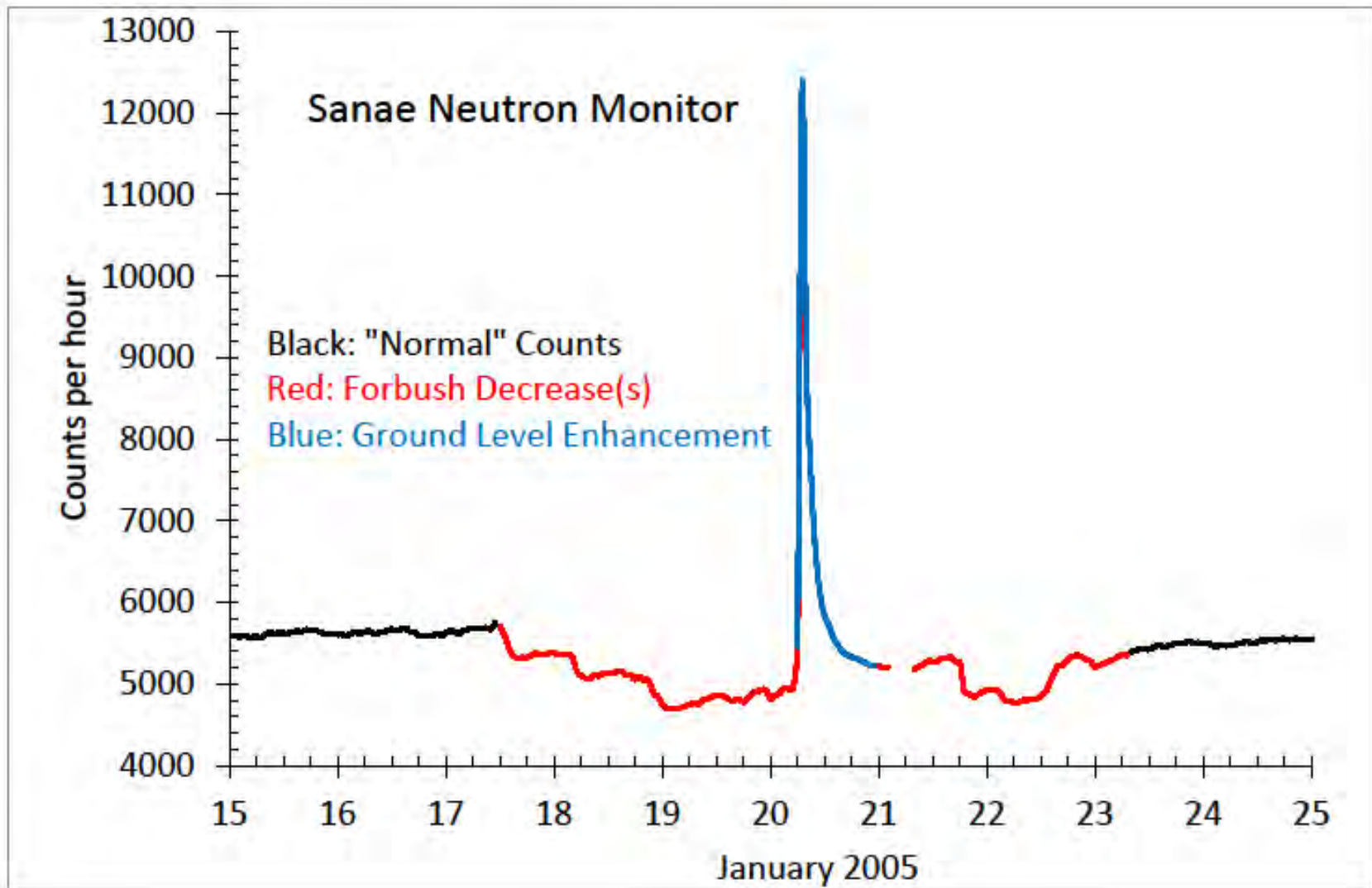
Solar Flare



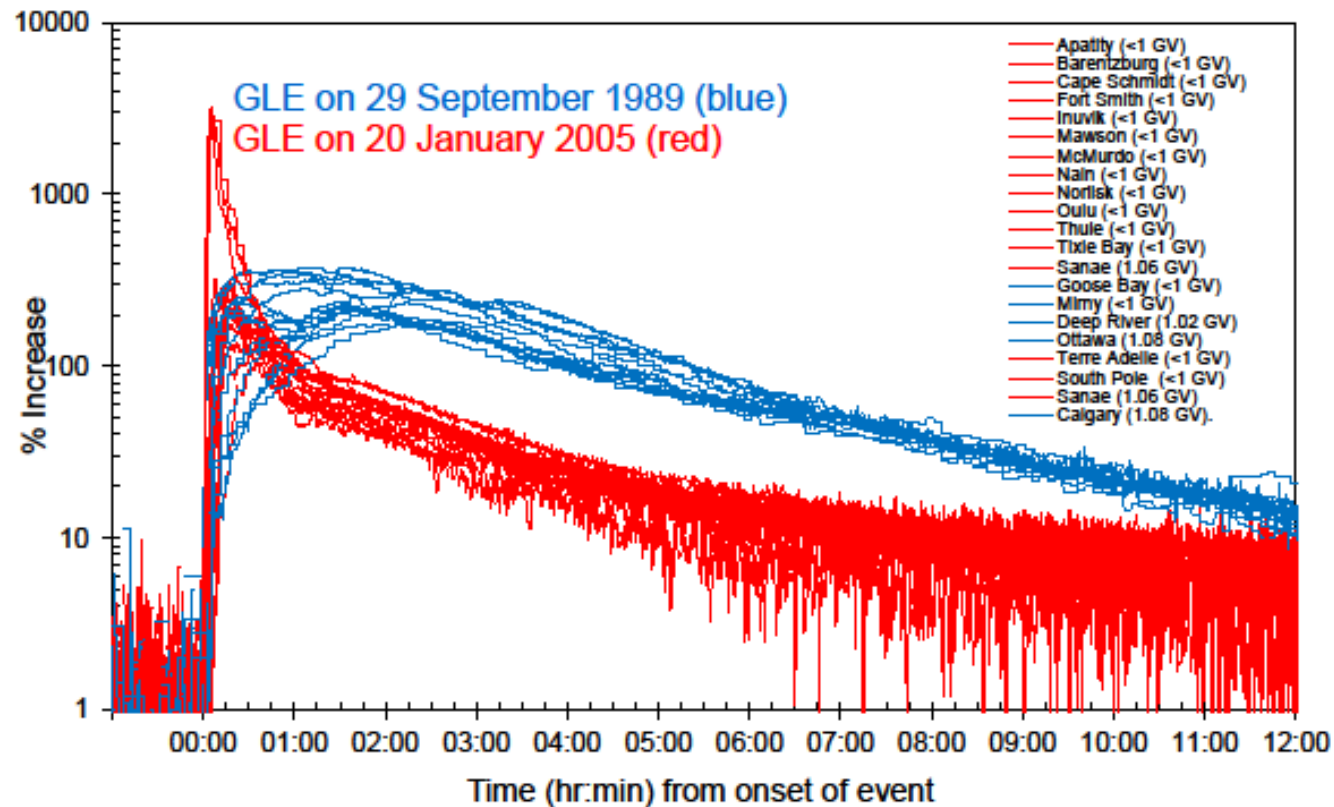
Ground-level Enhancement (GLE) = "cosmic" rays from sun



Ground-level ~~Enhancement~~ Excitement (GLE)



Two real big ones



Why?

- **Because it is there..... (academic)**
- **Appropriate Science and Technology**
- **Contribution to climate change**
- **Interdisciplinary (M.Sc. and M.Eng.)**

Three forms of the Transport Equation

$$\frac{\partial U}{\partial t} + \nabla \cdot (\mathbf{V}U - \mathbf{K} \cdot \nabla U) - \frac{1}{3} (\nabla \cdot \mathbf{V}) \frac{\partial}{\partial p} (pU) = 0$$

or, in terms of f

$$\frac{\partial f}{\partial t} + \nabla \cdot (\mathbf{V}f - \mathbf{K} \cdot \nabla f) - \frac{1}{3p^2} (\nabla \cdot \mathbf{V}) \frac{\partial}{\partial p} (p^3 f) = 0$$

or, slightly manipulated

$$\frac{\partial f}{\partial t} + \mathbf{V} \cdot \nabla f - \nabla \cdot (\mathbf{K} \cdot \nabla f) - \frac{1}{3p^2} (\nabla \cdot \mathbf{V}) \frac{\partial f}{\partial \ln p} = 0$$

Not changed for 45 years

Too difficult to solve analytically

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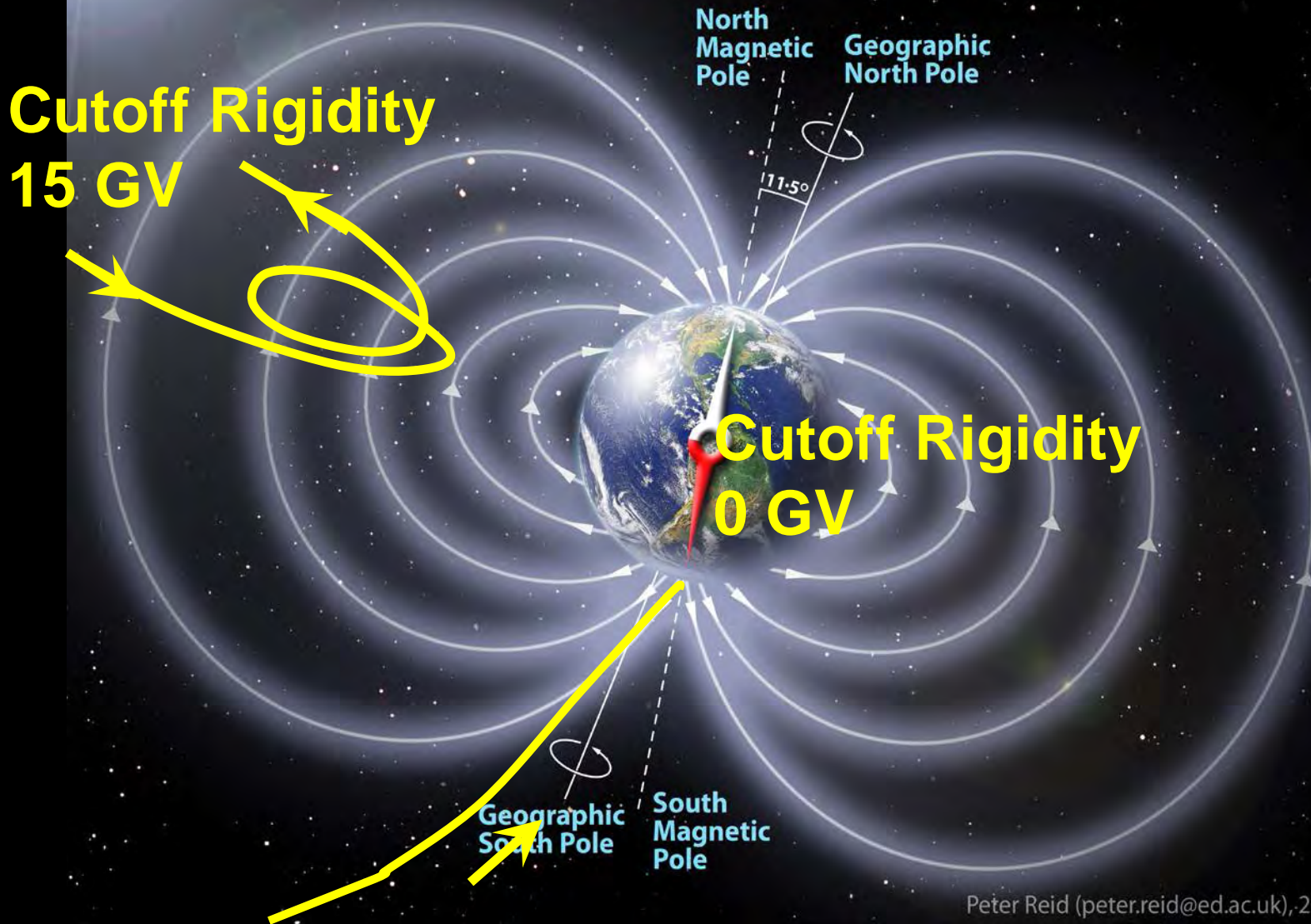
Pre-2006 SANAP Mission

To increase understanding of the **natural environment** and **life** in the Antarctic and Southern Ocean through appropriate science and technology

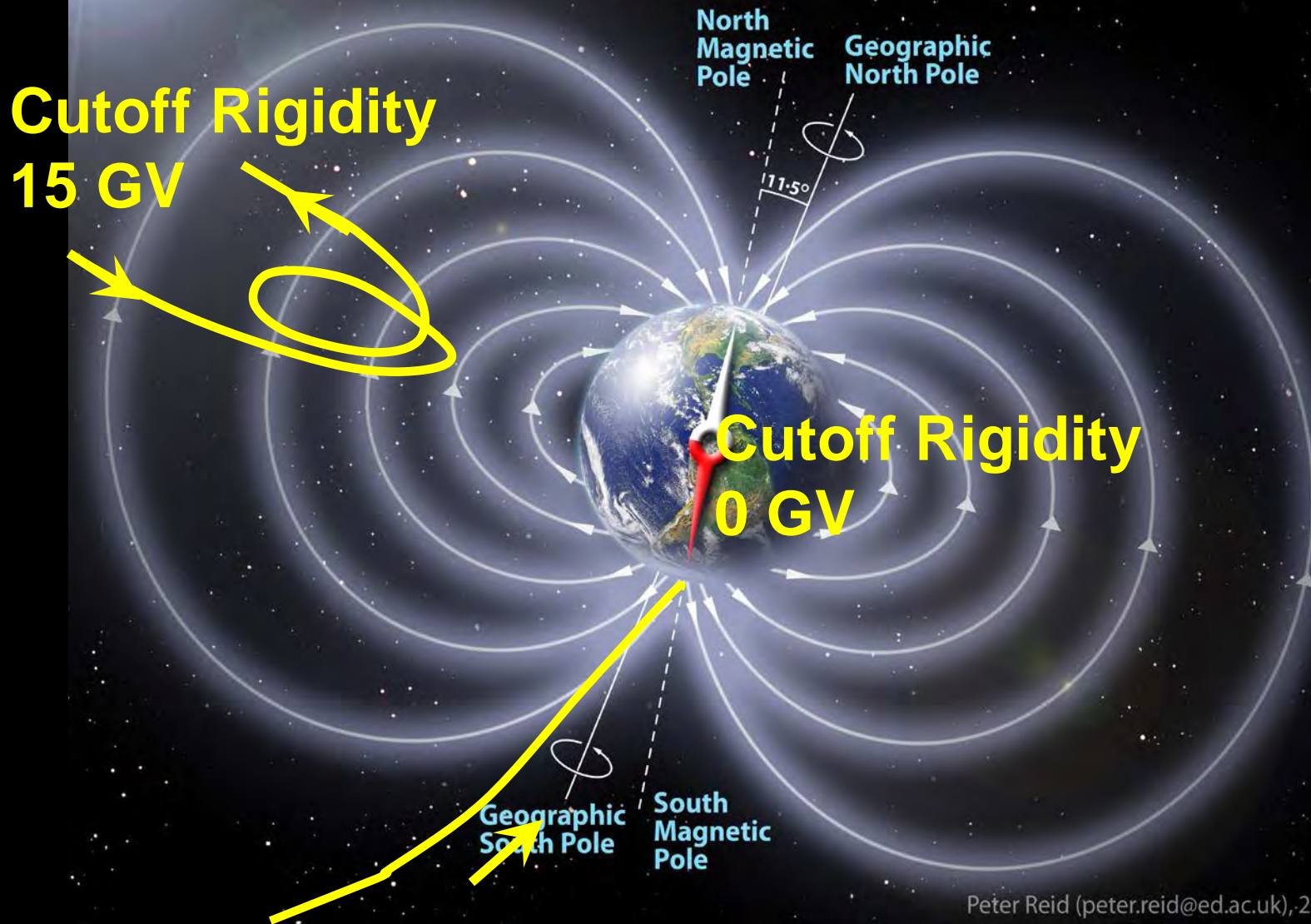
Pre-2006 SANAP Mission

To increase understanding of the **natural environment** and **life** in the Antarctic and Southern Ocean through **APPROPRIATE** science and technology

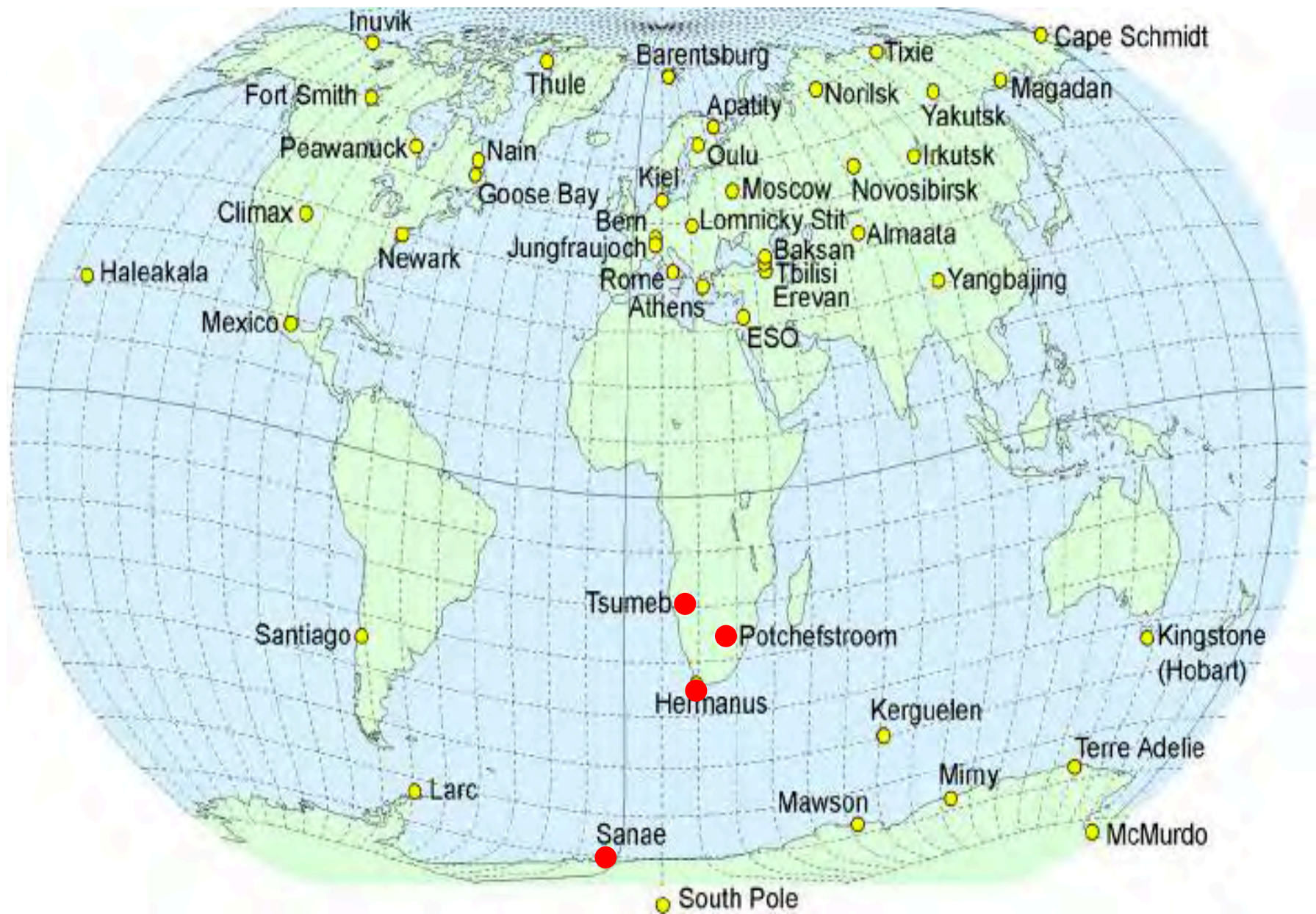
The poles are better



The poles are better – a window into geospace



Appropriate.....



Appropriate.....

- **One space mission > \$ 70 M = R 700 M**

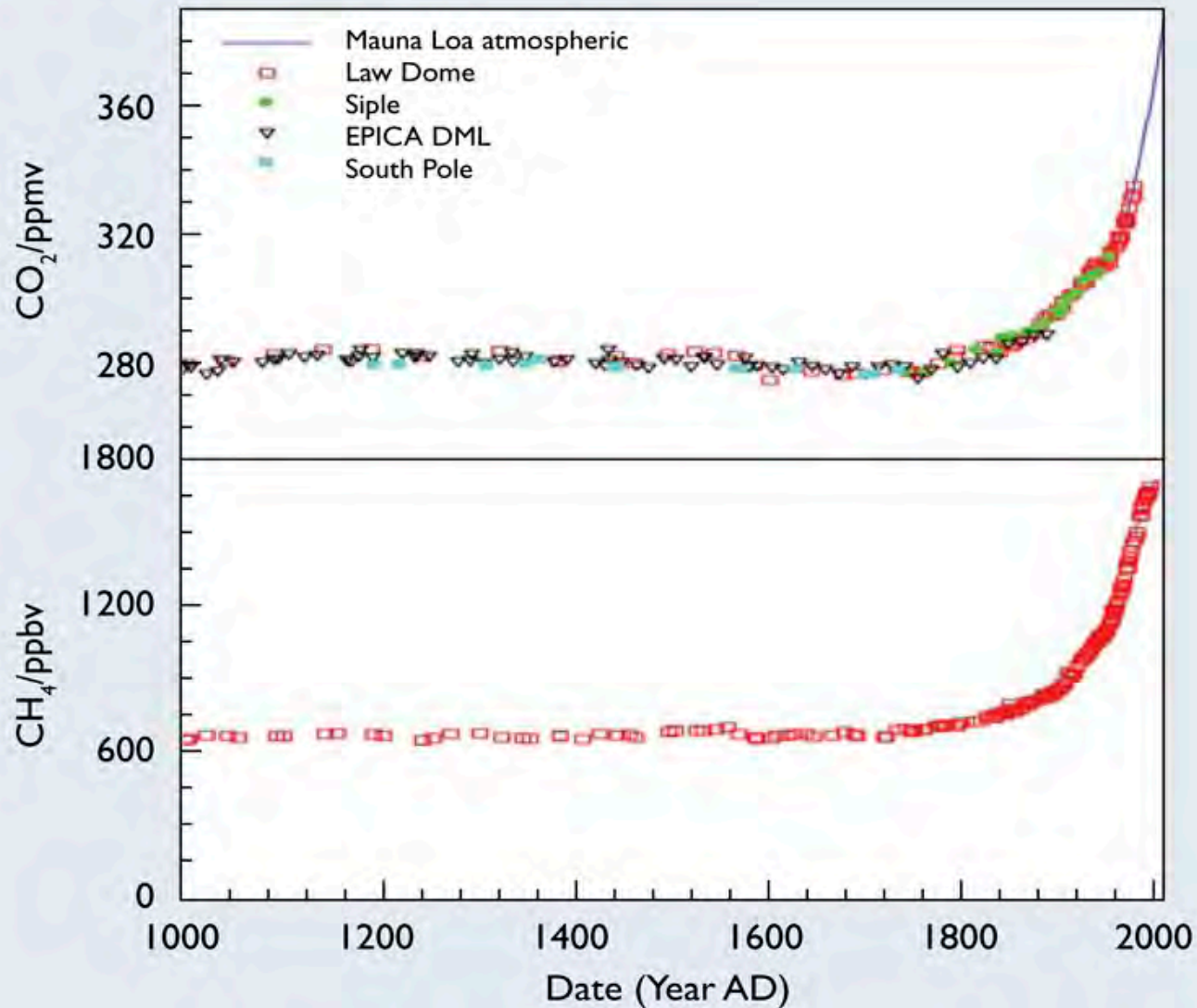
- **Neutron monitors = R 250 K per year x 40 NMs**
x 60 years = R 600 M

Why?

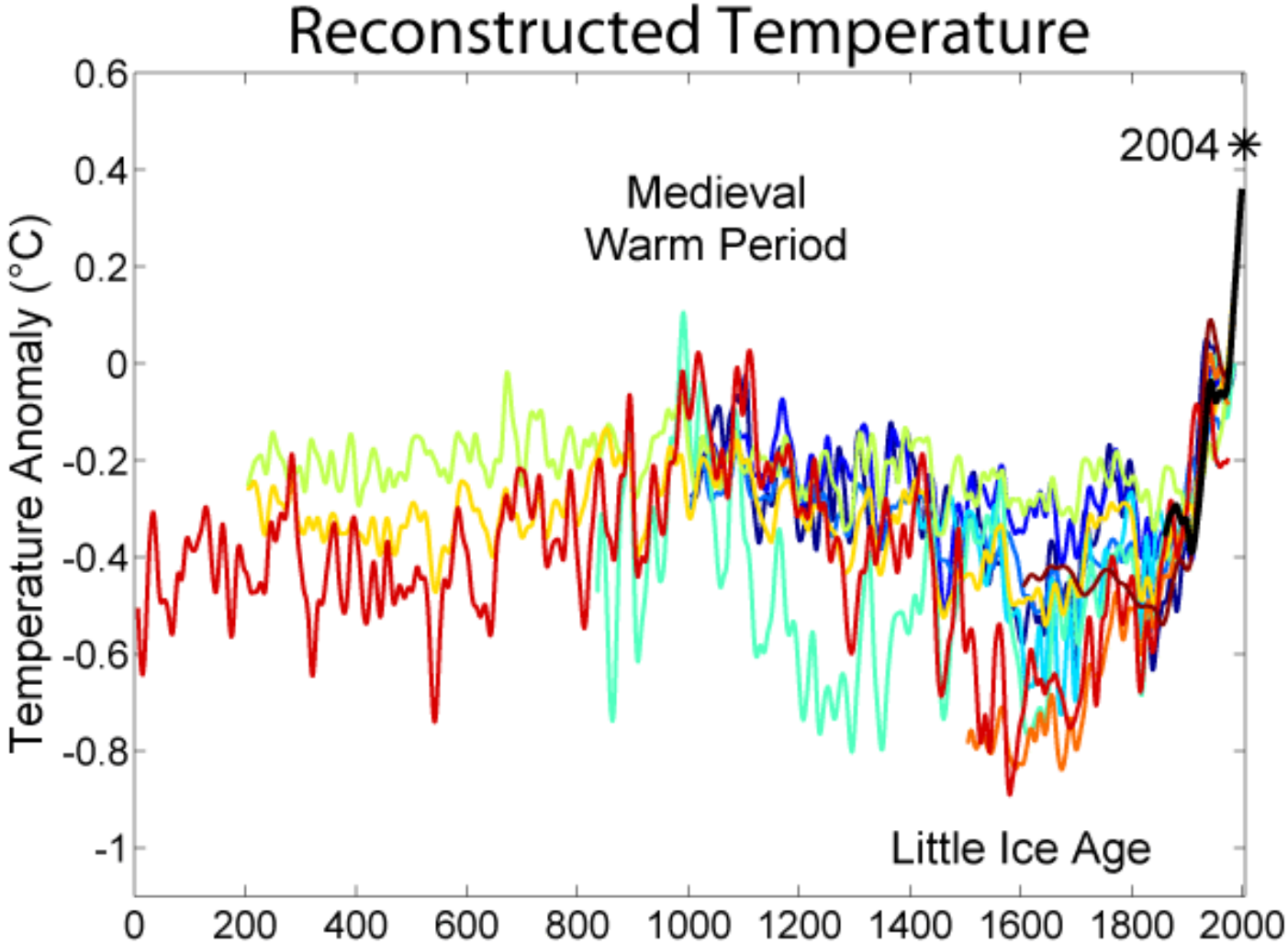
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Climate change

Fig. 2: CO_2 and CH_4 over the last 1,000 years⁽¹⁻⁴⁾

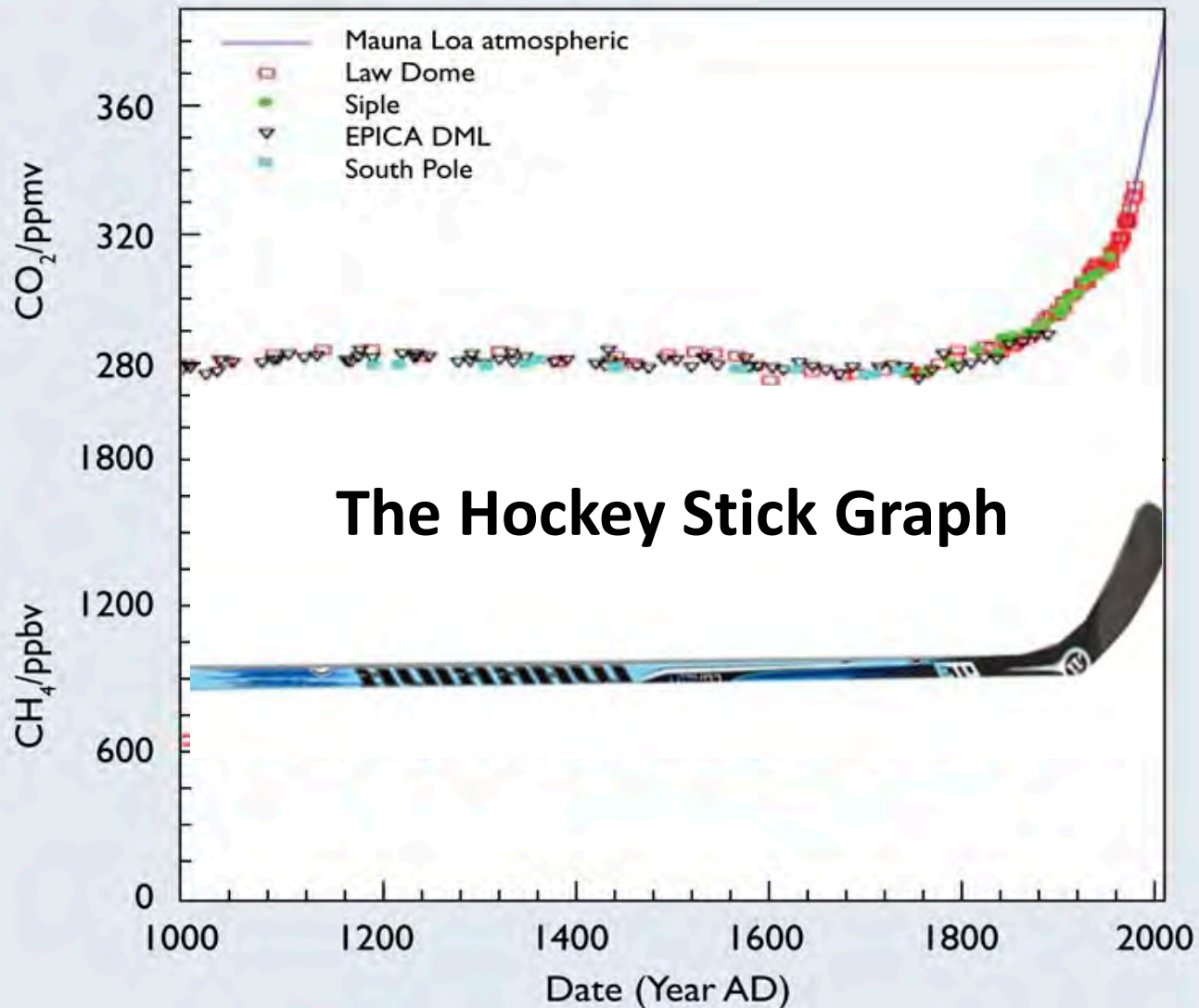


Climate change



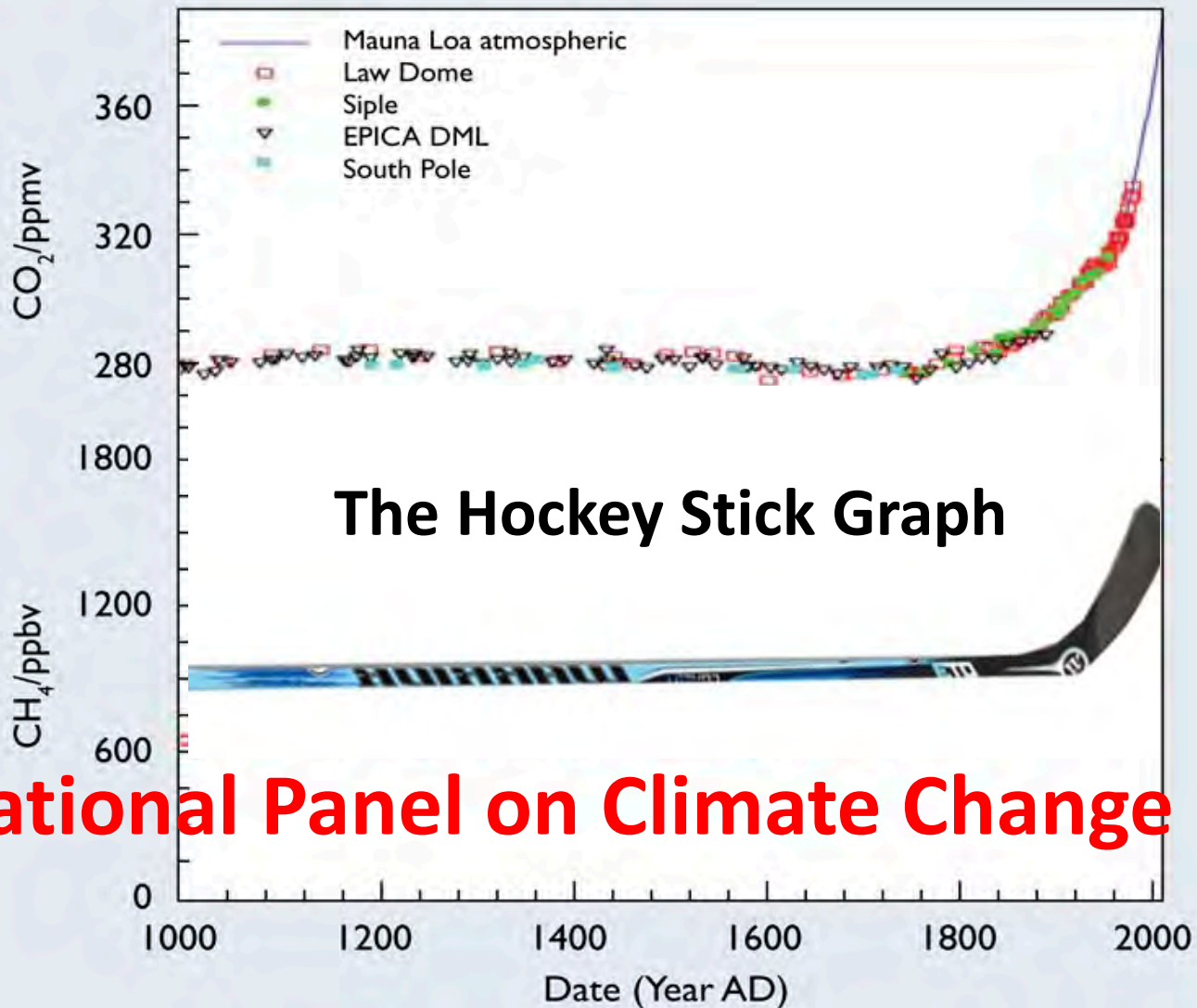
Climate change

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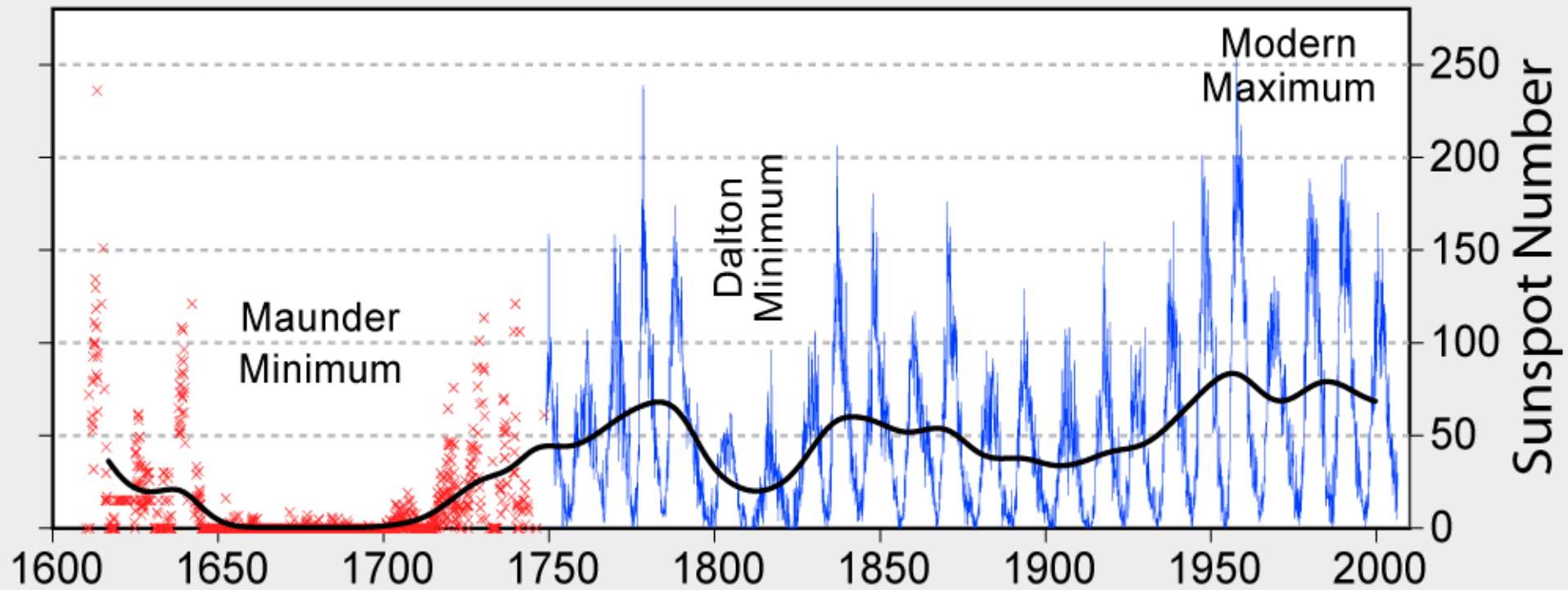
Climate change

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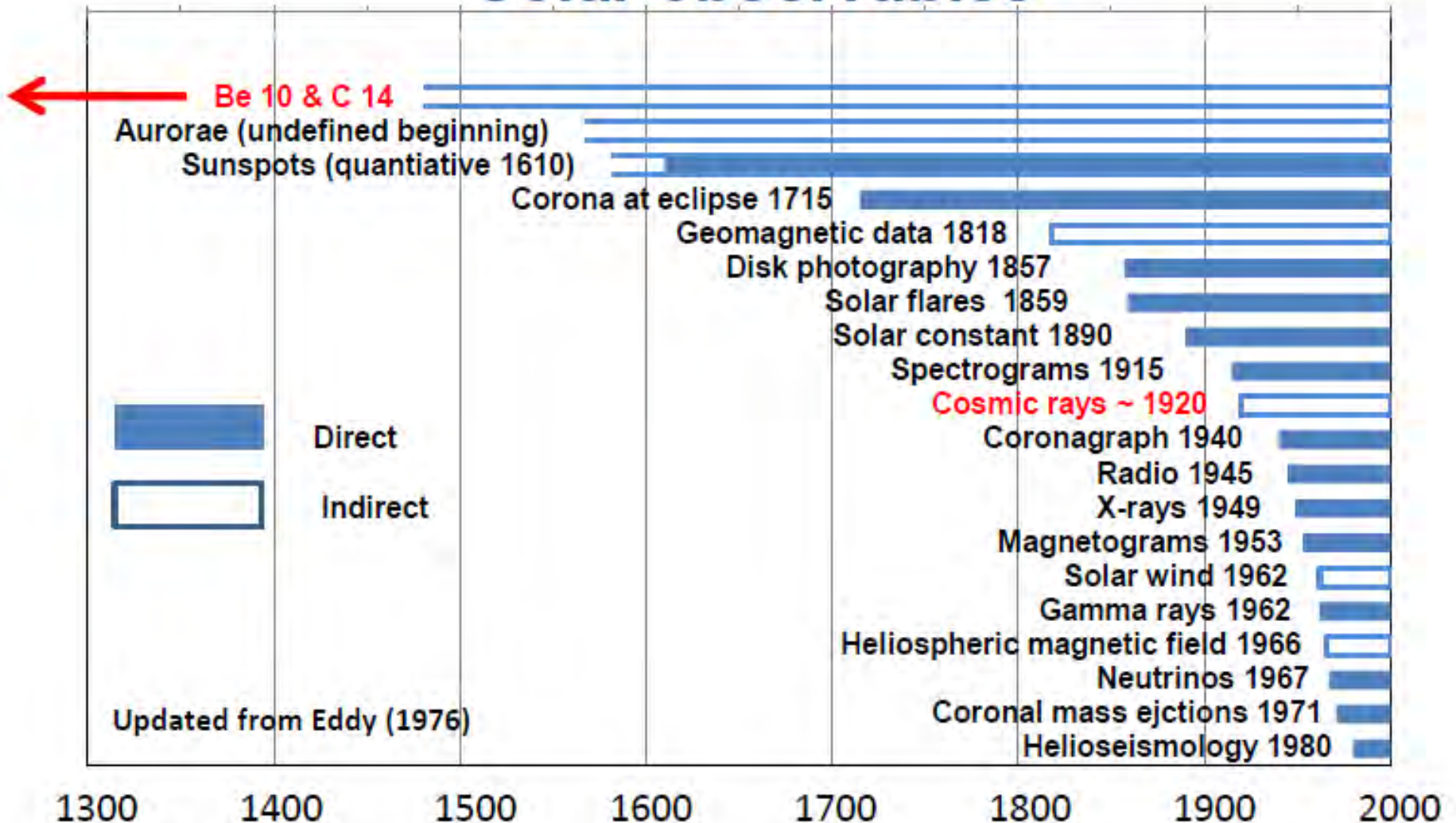
Climate change

400 Years of Sunspot Observations

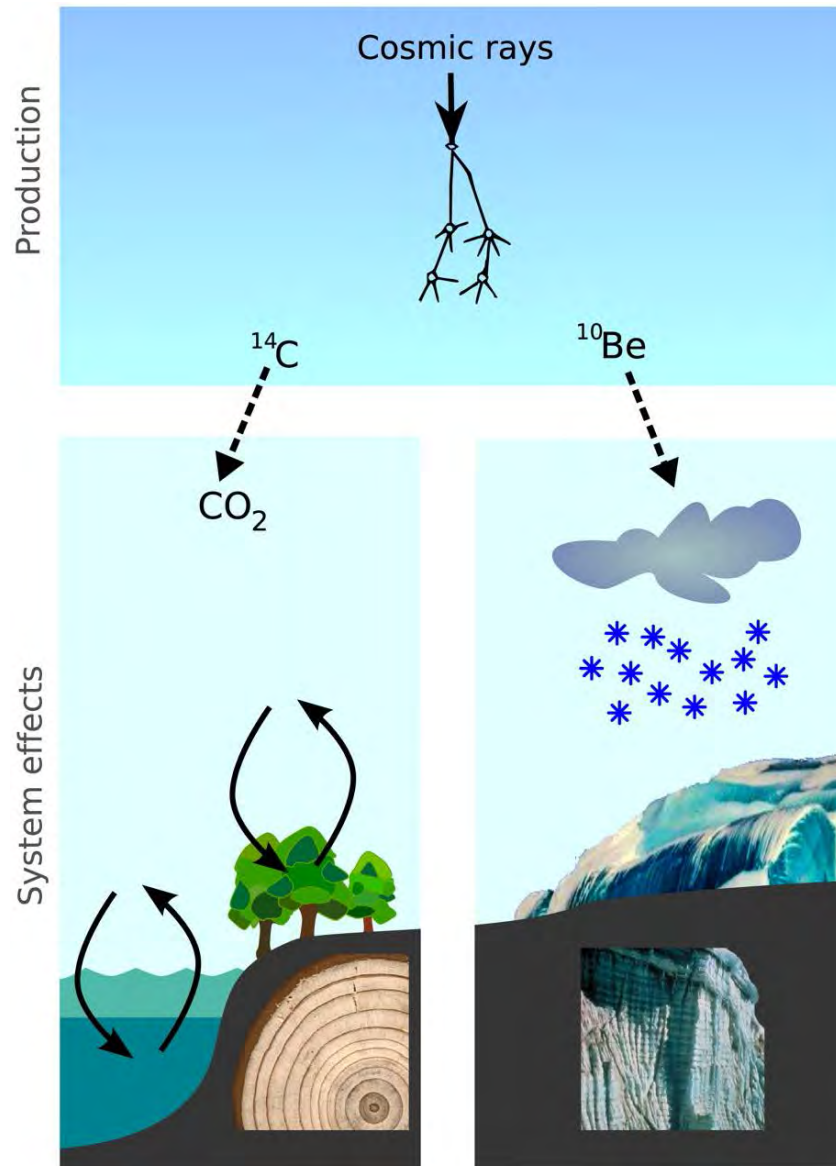


History of the Sun

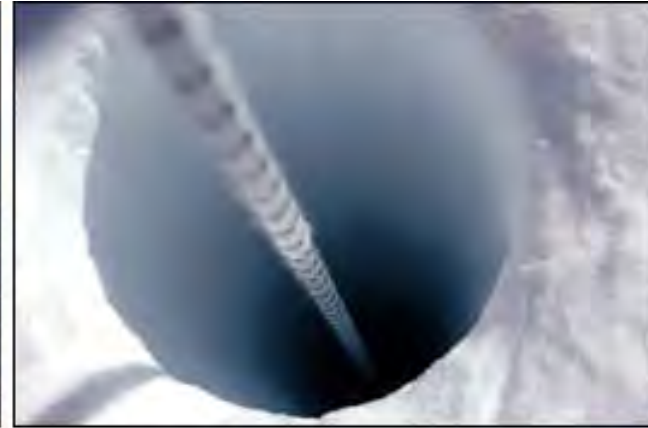
Solar observables



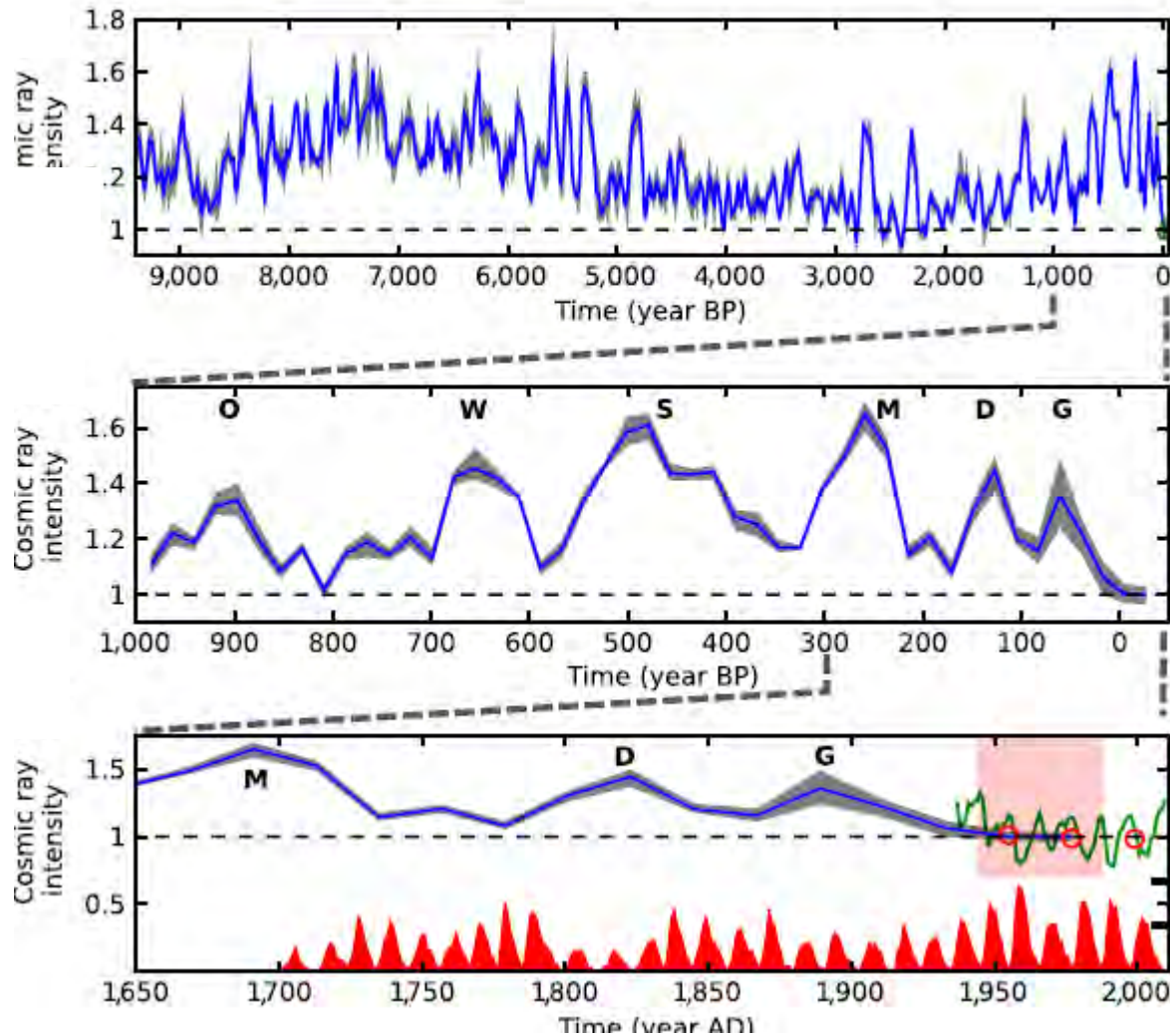
Paleo-cosmic rays: ^{14}C and ^{10}Be



Paleo-cosmic rays: ^{10}Be

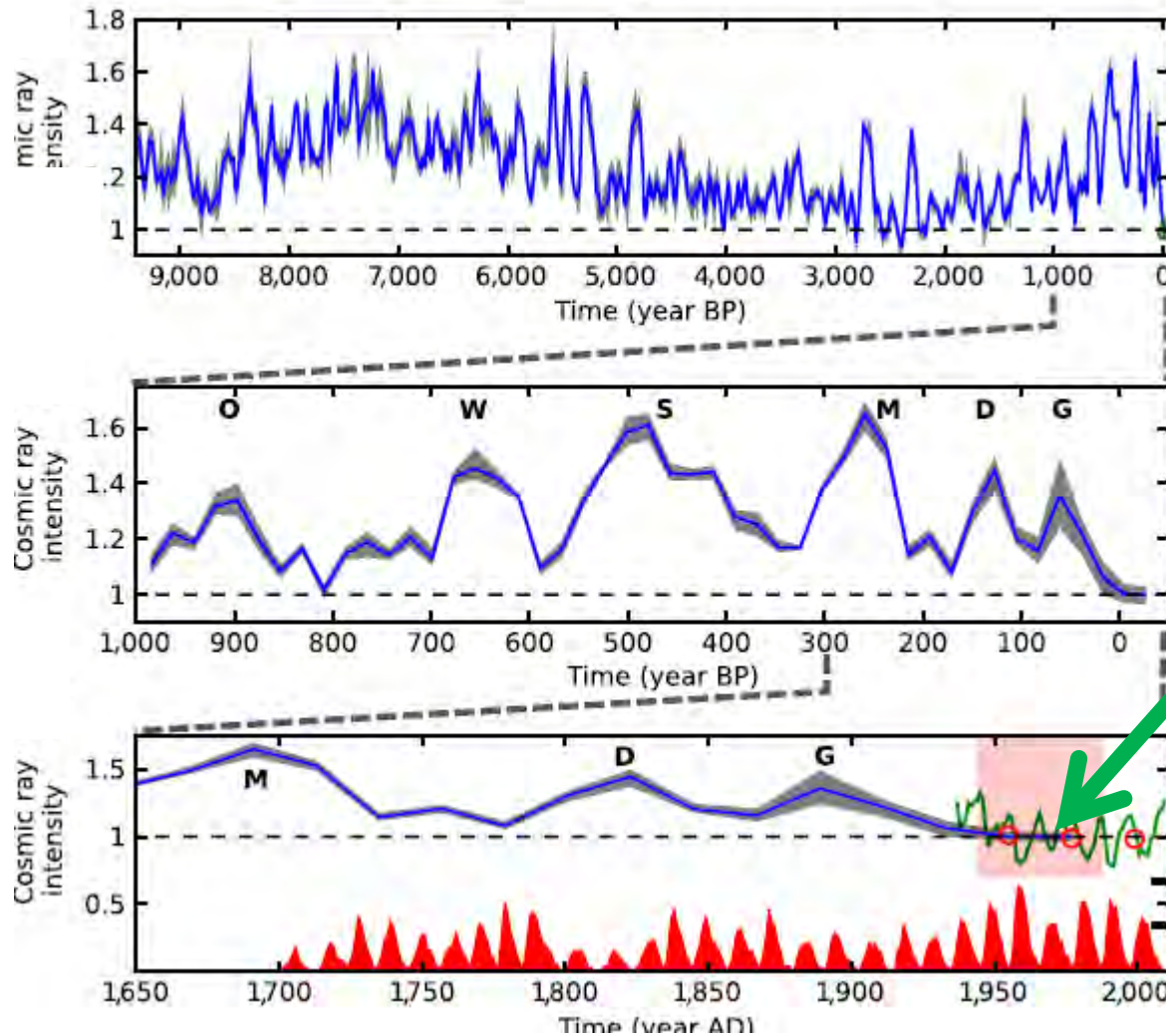


^{10}Be in polar ice ... Earth's neutron monitor



Steinhilber et al. (2012)

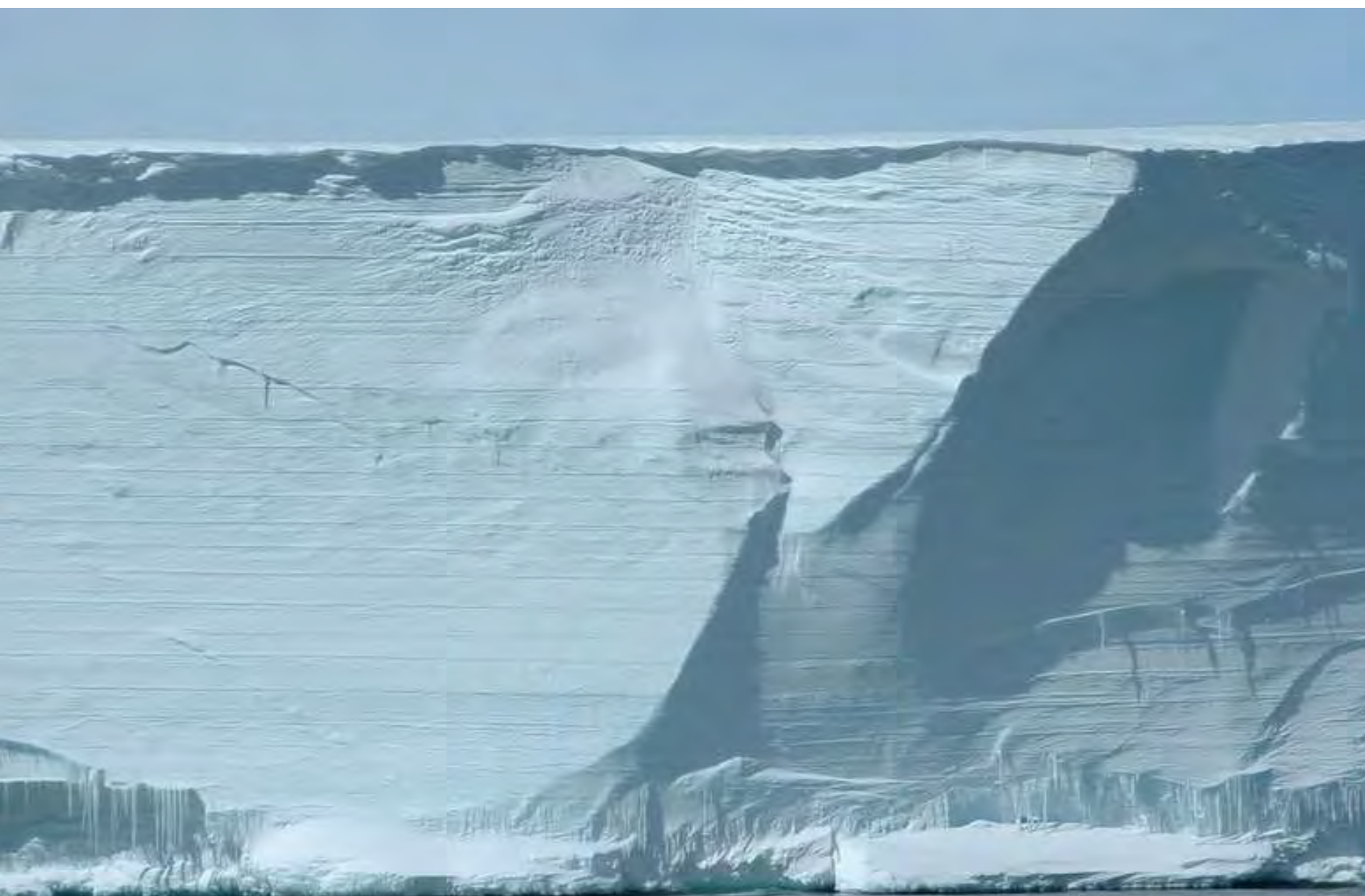
^{10}Be in polar ice ... Earth's neutron monitor



Real
neutron
monitor

Steinhilber et al. (2012)

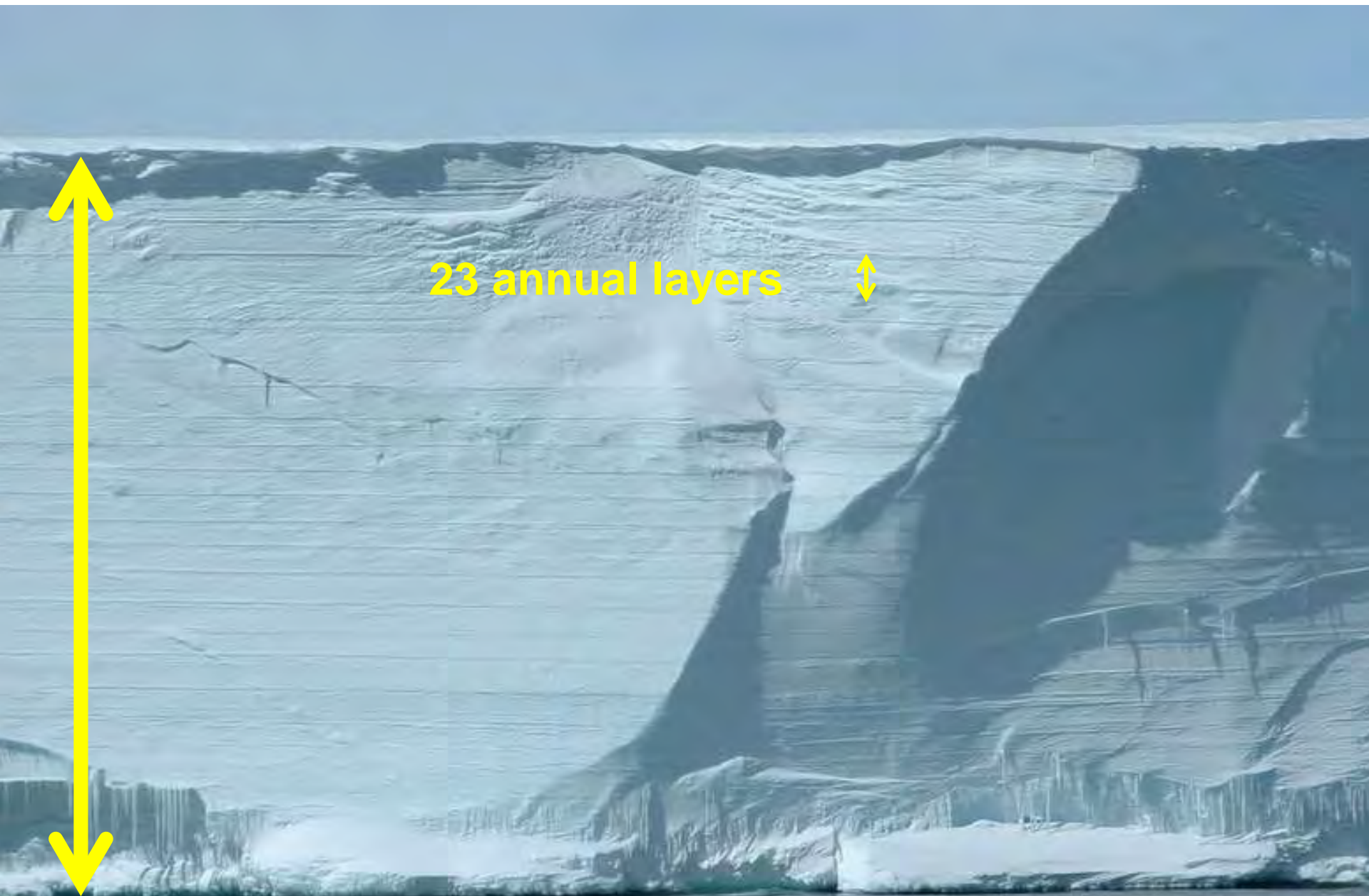
Ice shelf in Queen Maud Land



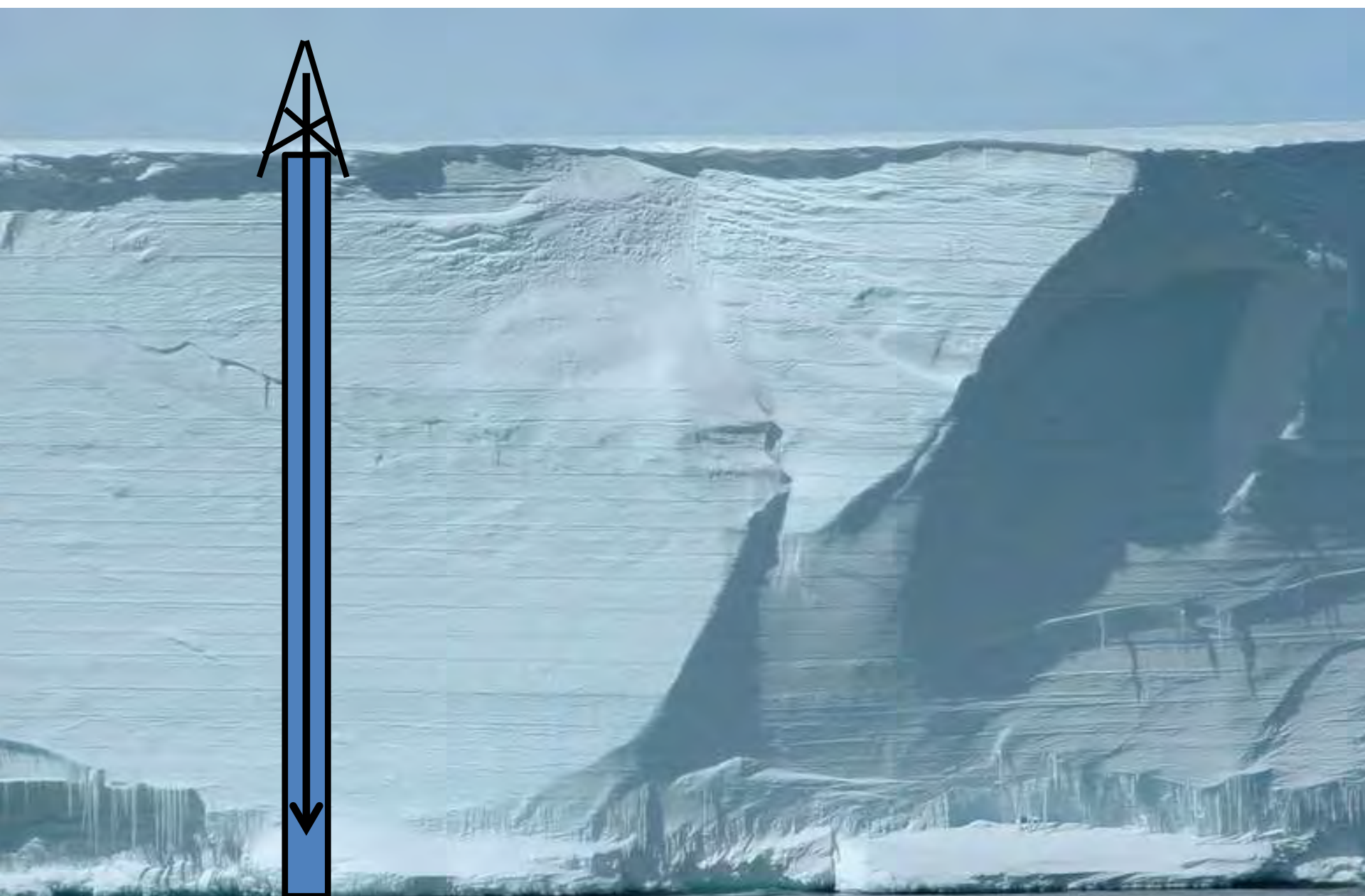
Ice shelf in Queen Maud Land



Ice shelf in Queen Maud Land



Shallow drilling 60 years deep

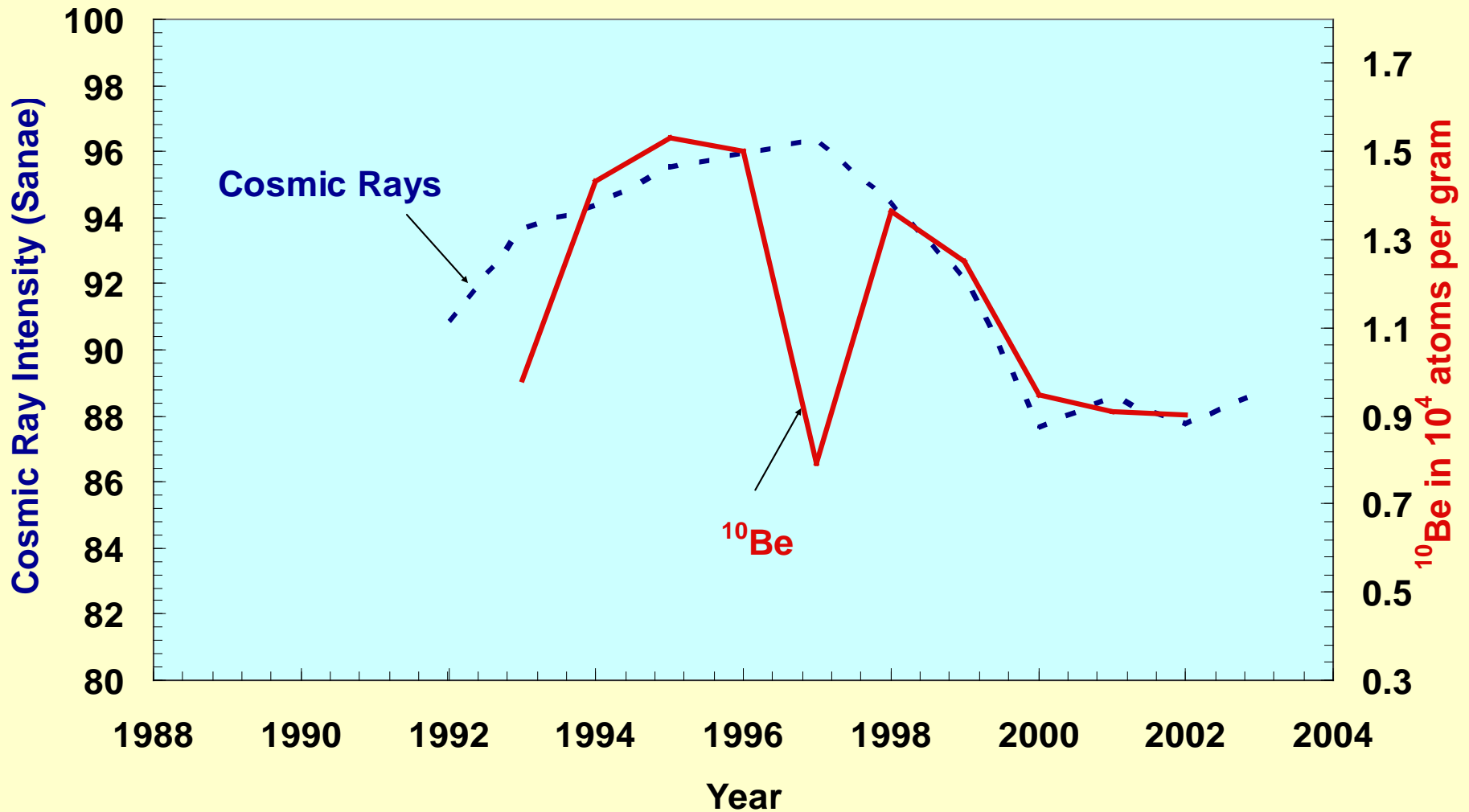


Pilot Project 2006



Pilot project 2006

^{10}Be and Cosmic Rays



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Thank you:

NRF/DST

DEA

Rhodes University Organisers

Cosmic-ray spectrum

